

# First Results for Electron Neutrino Appearance in MINOS

Lisa Whitehead  
BNL

BNL HEP Seminar  
March 9, 2009



# Outline

- ♦ Brief intro to neutrino oscillation
- ♦ MINOS beam and detectors
- ♦ The  $\nu_e$  appearance analysis:
  - ♦ selection of candidate  $\nu_e$  events
  - ♦ background analysis with near detector data
  - ♦ far detector prediction
  - ♦ results
- ♦ Summary

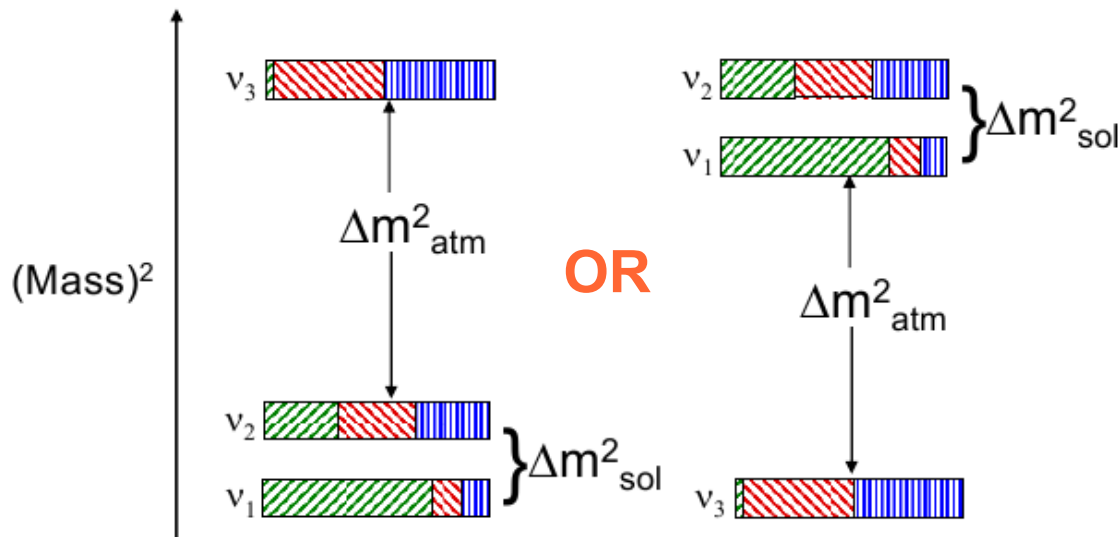
# Neutrino Oscillation

$$S_{jk} = \sin \theta_{jk}$$

$$C_{jk} = \cos \theta_{jk}$$

$$U_{MNSP} = \underbrace{\begin{pmatrix} 1 & 0 & 0 \\ 0 & C_{23} & S_{23} \\ 0 & -S_{23} & C_{23} \end{pmatrix}}_{\text{atmospheric}} \underbrace{\begin{pmatrix} C_{13} & 0 & S_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -S_{13}e^{i\delta} & 0 & C_{13} \end{pmatrix}}_{\delta \text{ and } \theta_{13} \text{ are unknown}} \underbrace{\begin{pmatrix} C_{12} & S_{12} & 0 \\ -S_{12} & C_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}}_{\text{solar}}$$

$$\Delta m_{ij}^2 = m_i^2 - m_j^2$$



Mass hierarchy is unknown

Best limit by CHOOZ reactor  
neutrino experiment:  
 $\sin^2 2\theta_{13} < 0.15$

For the remainder of this talk,  
“signal” plots and numbers  
assume:

$$\Delta m_{31}^2 \sim \Delta m_{32}^2 > 0$$

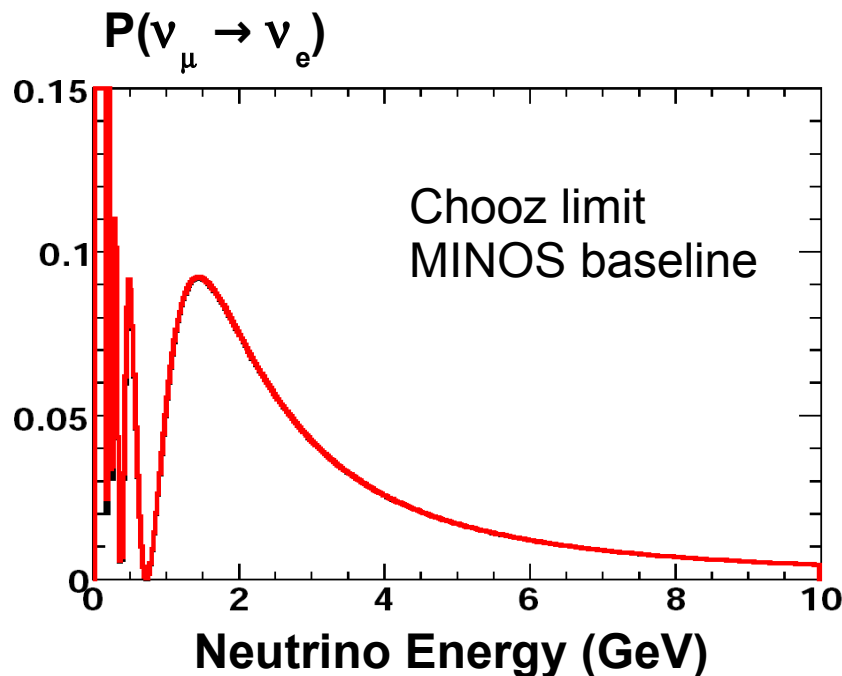
$$\sin^2 2\theta_{13} = 0.15$$

$$\delta = 0$$

# Measuring $\theta_{13}$

$\nu_e$  appearance  
probability

$$P(\nu_\mu \rightarrow \nu_e) \approx \sin^2 2\theta_{13} \sin^2 \theta_{23} \frac{\sin^2(A-1)\Delta}{(A-1)^2} \\ + 2\alpha \sin \theta_{13} \cos \delta \sin 2\theta_{12} \sin 2\theta_{23} \frac{\sin A\Delta}{A} \frac{\sin(A-1)\Delta}{(A-1)} \cos \Delta \\ - 2\alpha \sin \theta_{13} \sin \delta \sin 2\theta_{12} \sin 2\theta_{23} \frac{\sin A\Delta}{A} \frac{\sin(A-1)\Delta}{(A-1)} \sin \Delta \\ + \alpha^2 \sin^2 2\theta_{12} \cos^2 \theta_{23} \frac{\sin^2 A\Delta}{A^2}$$



**Matter effect:**

$$A \equiv \frac{G_f n_e L}{\sqrt{2}\Delta} \approx \frac{E}{11 \text{ GeV}}$$

$$\Delta \equiv \frac{\Delta m_{31}^2 L}{4E}$$

$$\alpha \equiv \frac{\Delta m_{21}^2}{\Delta m_{31}^2}$$

$\theta_{13}$  is coupled with  $\delta$  and the mass hierarchy

# MINOS (Main Injector Neutrino Oscillation Search)

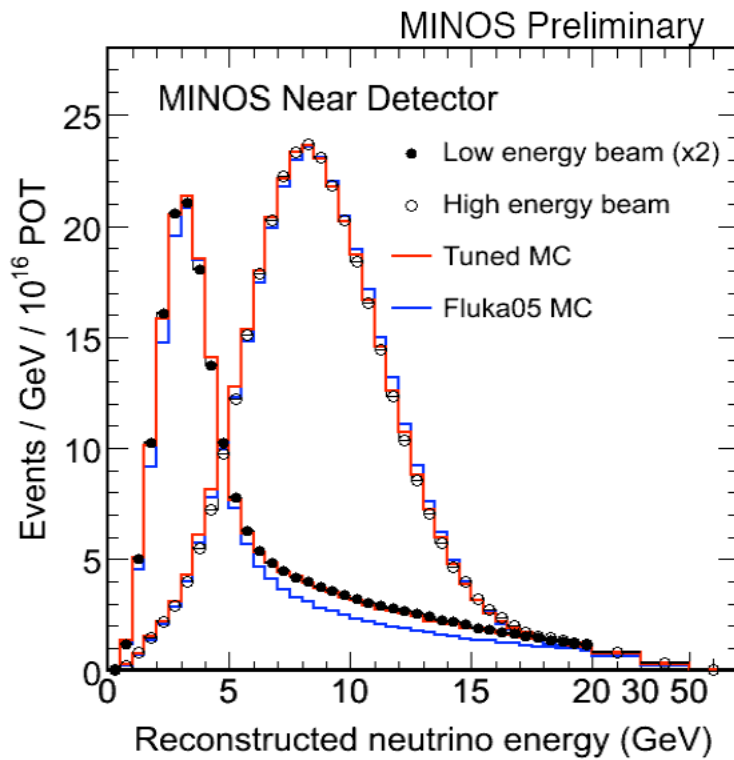
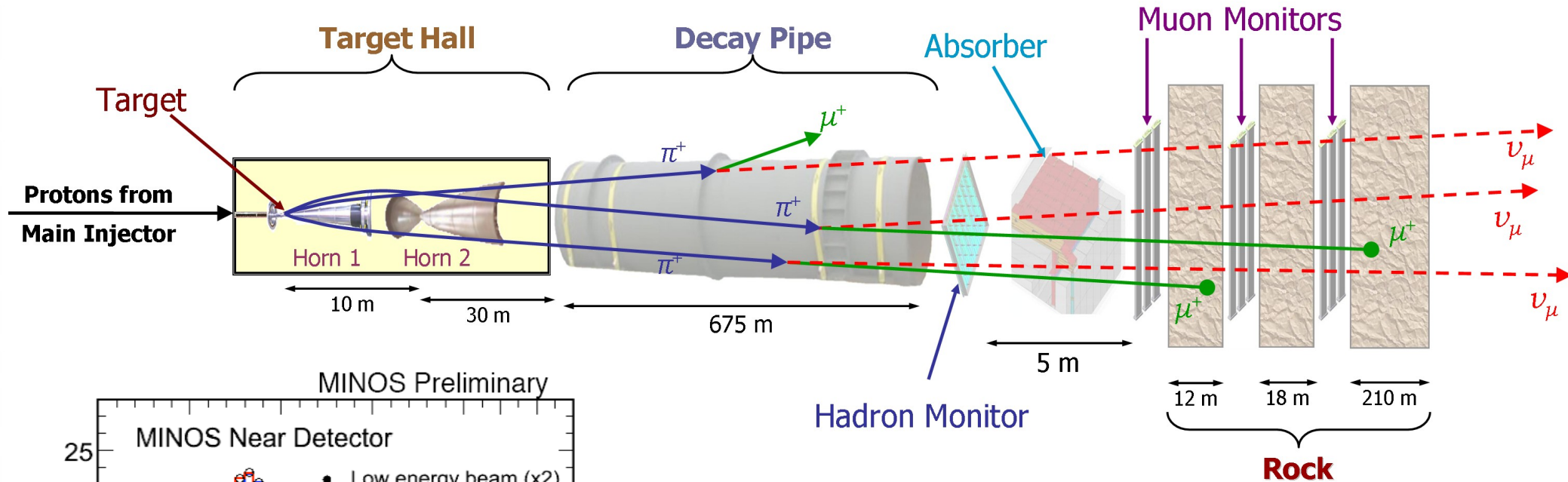
Produce a beam of muon neutrinos at Fermilab

Use data from a near detector to form a prediction for the far detector (number of events and/or energy spectrum)

Neutrino oscillation will cause a deviation from the prediction in the far detector

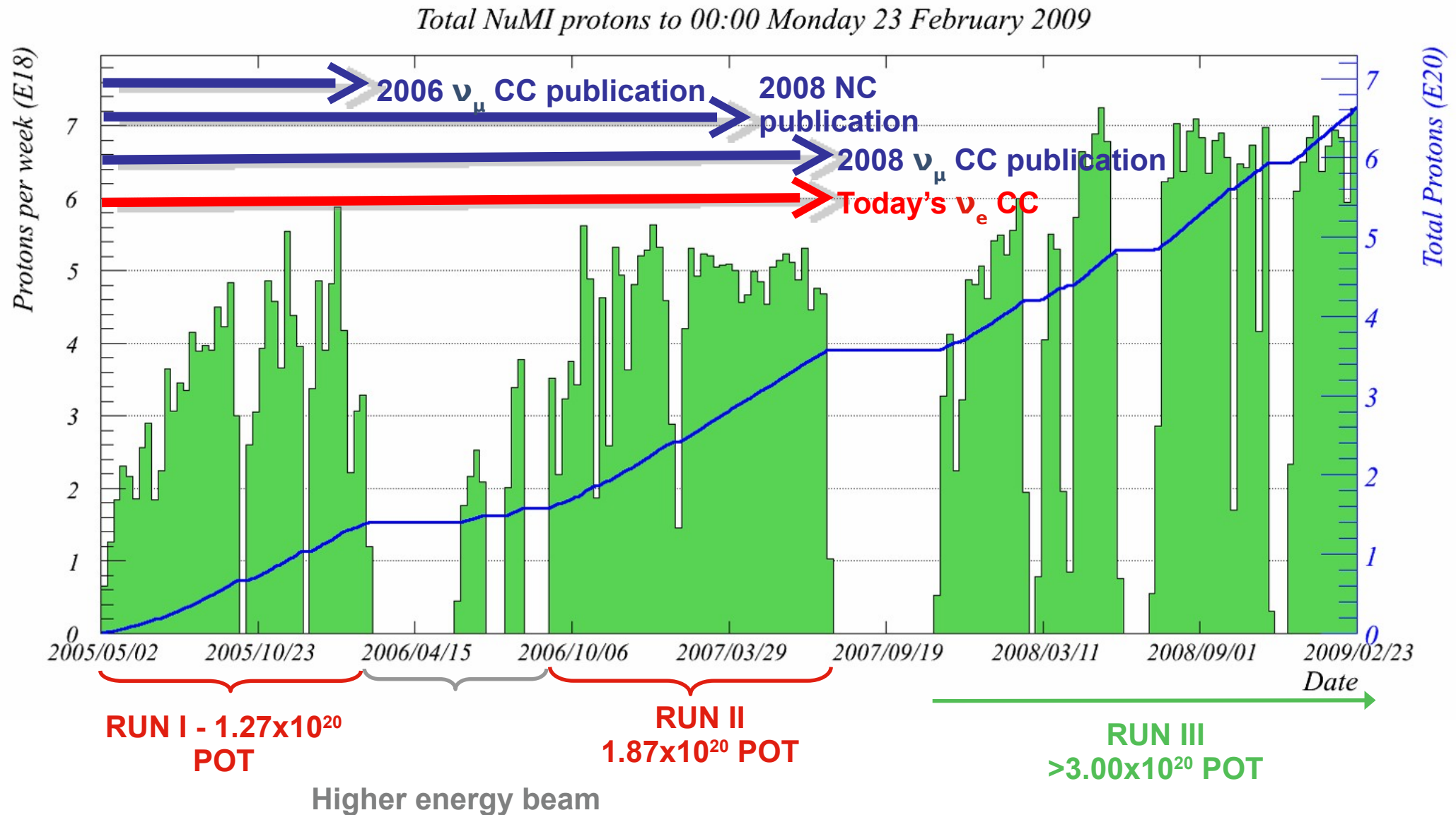


# NuMI Beam



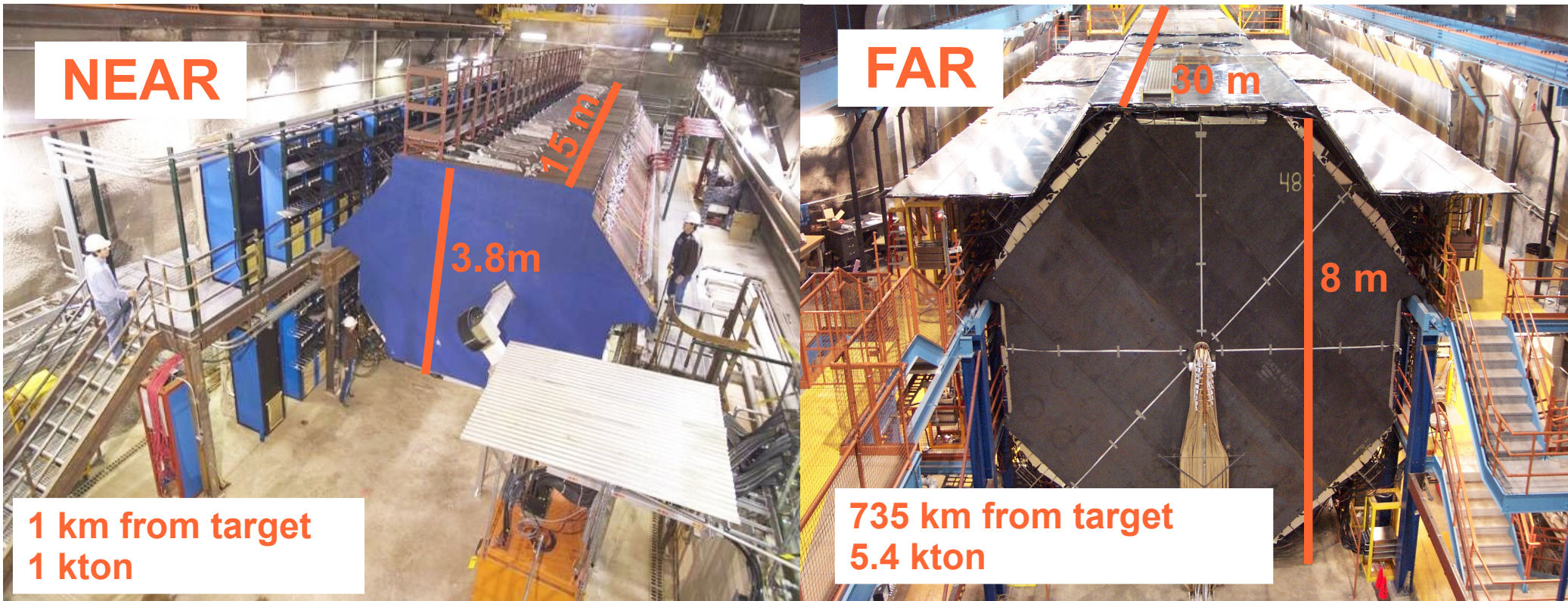
- 120 GeV protons from the Main Injector
- carbon target
- magnetic horns focus secondary particles
- particles decay to neutrinos
- target position affects energy spectrum
- timing structure: 10  $\mu$ s spill every 2.2 s
- currently,  $3 \times 10^{13}$  protons per spill
- Beam is mostly  $\nu_\mu$  with a small ( $\sim 1.3\%$ )  $\nu_e$  component which we know to 10%

# NuMI Beam: Protons-on-Target





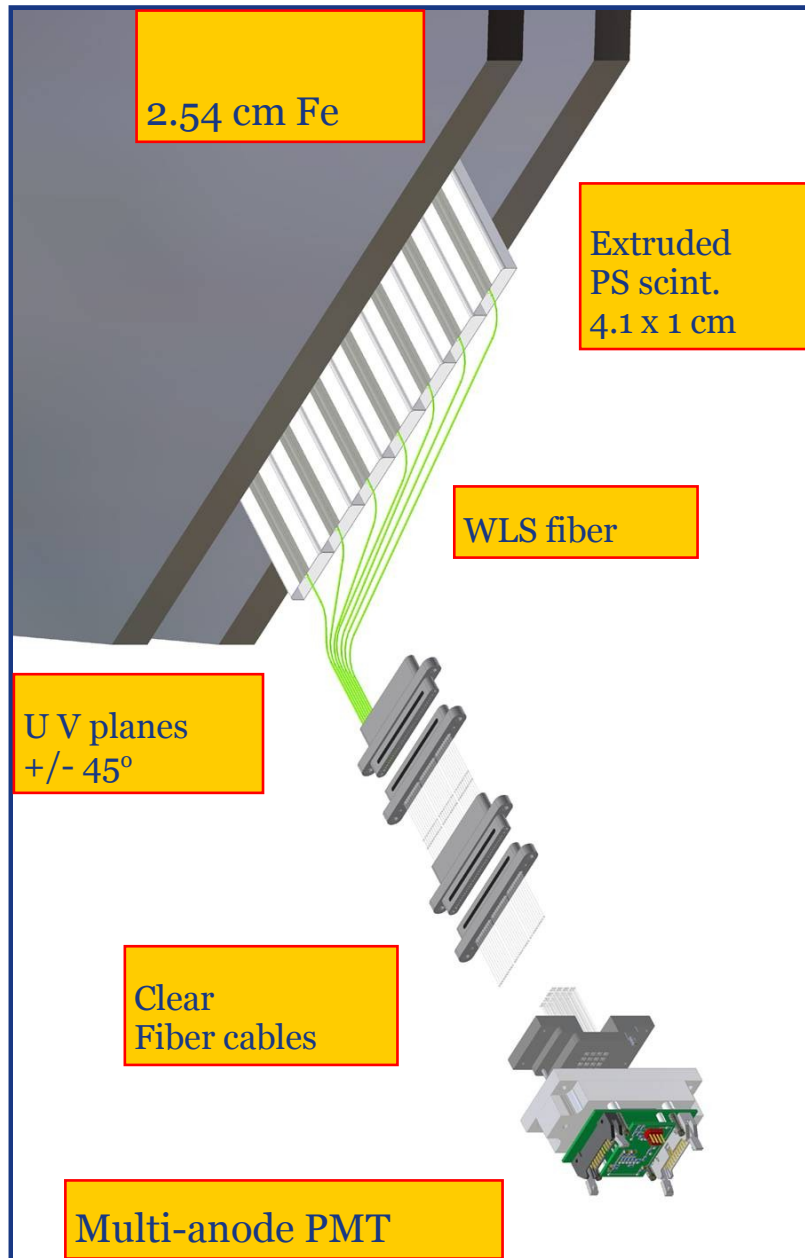
# MINOS detectors



alternating layers of steel plates and scintillator strips  
in a 1.3 T toroidal magnetic field



# MINOS detectors



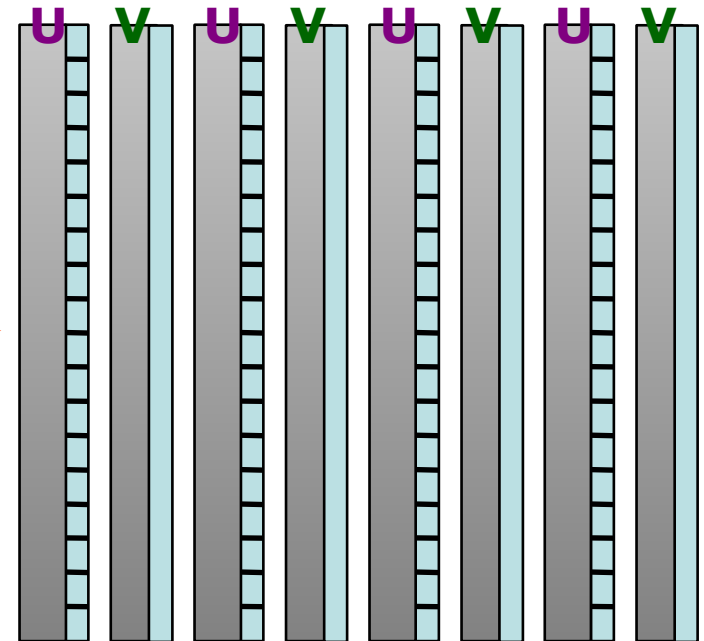
steel thickness: 2.54cm  $\sim 1.44X_0$   
strip width: 4.12cm (Moliere rad  $\sim 3.7$ cm)

Strips in adjacent planes are oriented orthogonally enabling 3D reconstruction

Strips have WLS fibers read out by multi-anode PMTs

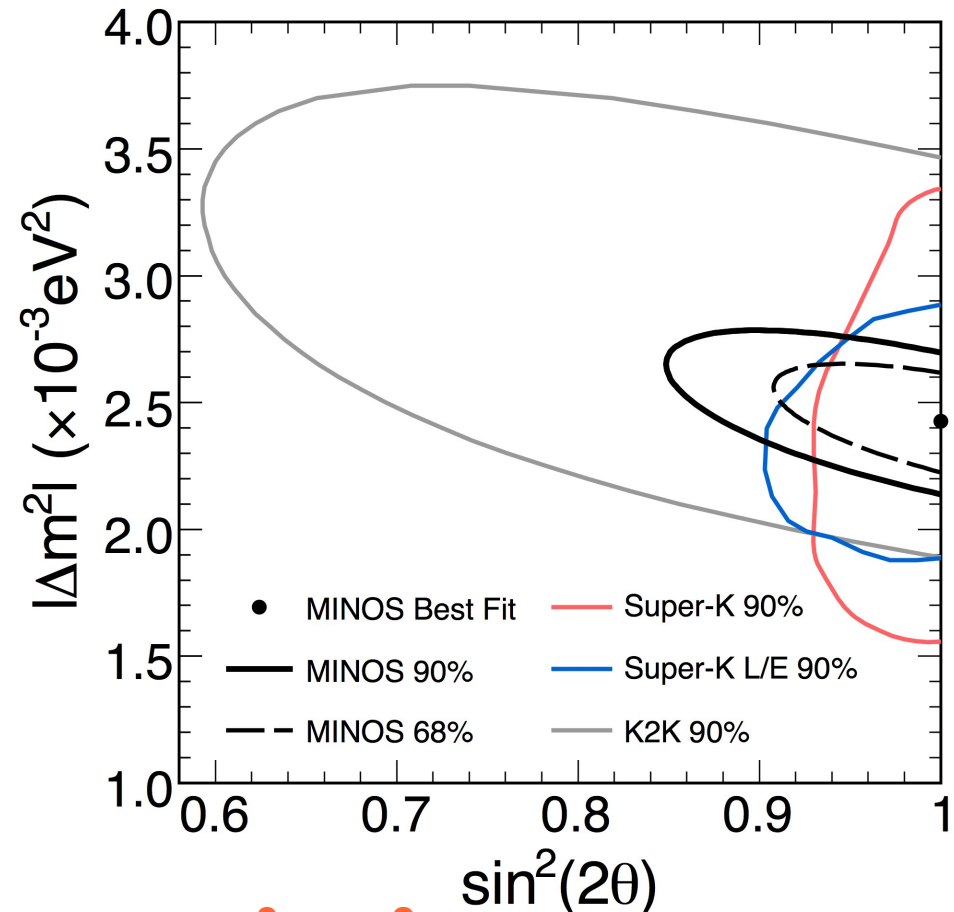
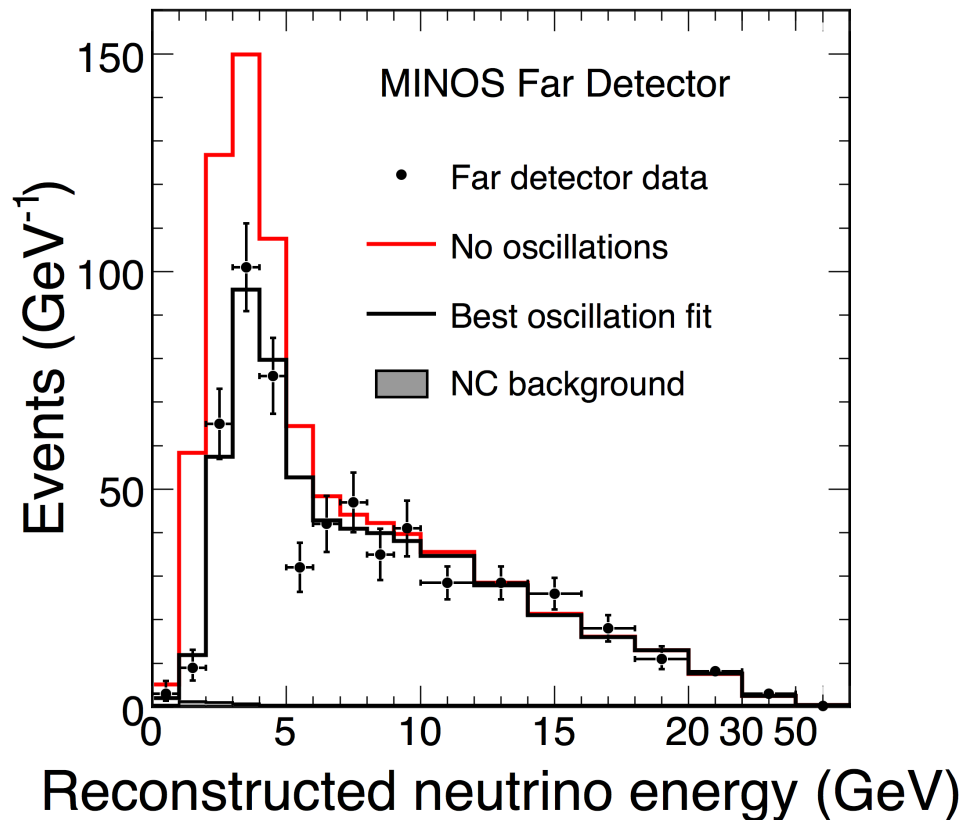
U/V strips  
oriented  $\pm 45^\circ$   
from vertical

beam →



# MINOS $\nu_\mu$ Disappearance

$$P(\nu_\mu \rightarrow \nu_\mu) = 1 - \sin^2(2\theta) \sin^2 \left( 1.27 \Delta m^2 \frac{L}{E} \right)$$



$$|\Delta m^2_{32}| = (2.43 \pm 0.13) \times 10^{-3} \text{eV}^2 \text{ (68\% CL)}$$

$$\sin^2 2\theta_{23} > 0.9 \text{ (90\% CL)}$$

Most precise measurement of  $|\Delta m^2_{32}|$  yet

# Searching for $\nu_e$ appearance

Determine the selection criteria for  $\nu_e$  candidate events.

Use near detector data (where we expect no oscillated  $\nu_e$ 's) to study the background.

Extrapolate the near detector background sample to get the far detector background prediction.

## Blind Analysis:

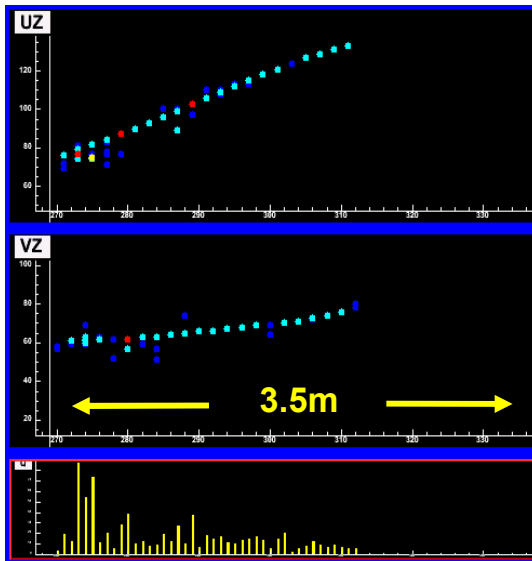
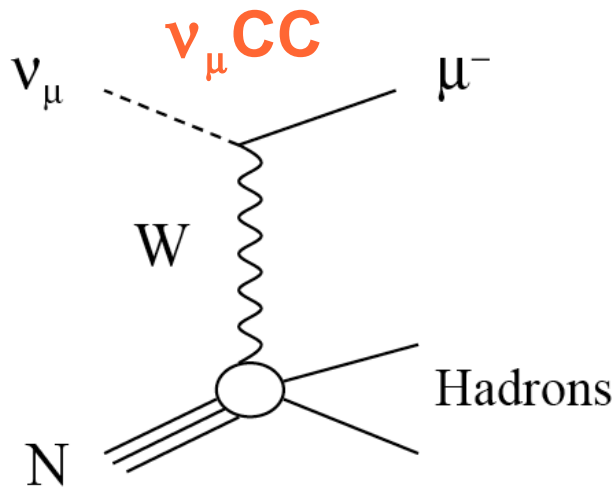
Sideband study - Examine far detector data outside of the signal region to test the analysis.

Open the box – Examine far detector data in the signal region and look for an excess of  $\nu_e$ -like events over predicted background in far detector.

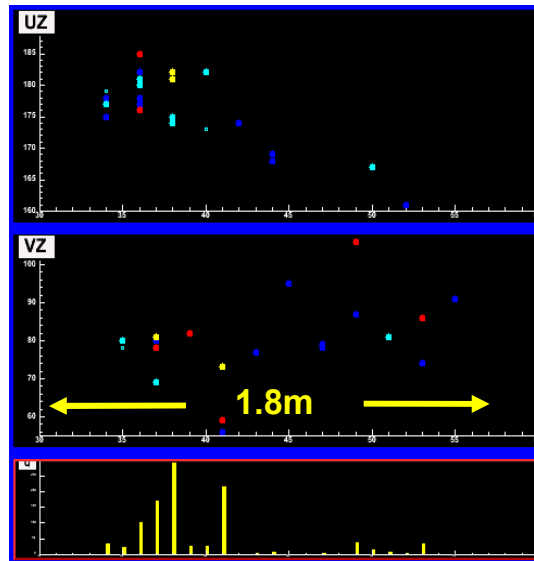
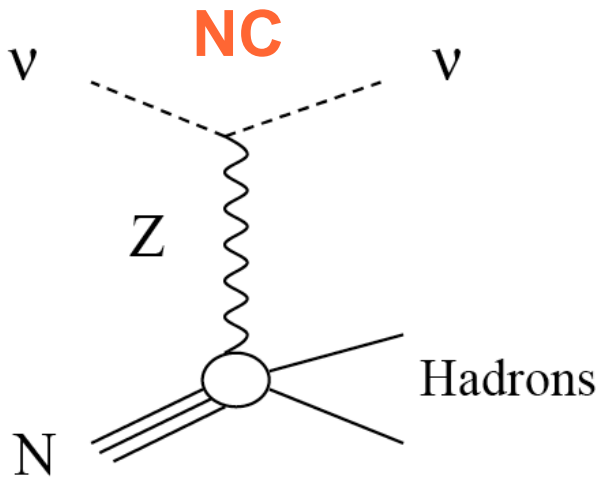
At the CHOOZ limit, we would expect an excess of only 6-12 events! (3.14e20 POT)

# Neutrino Events at MINOS

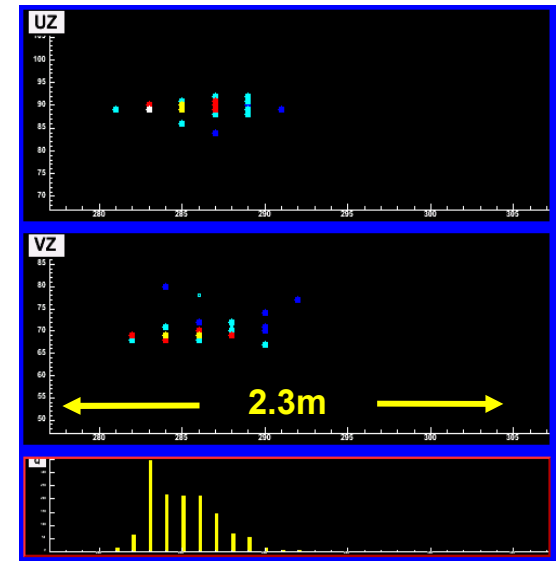
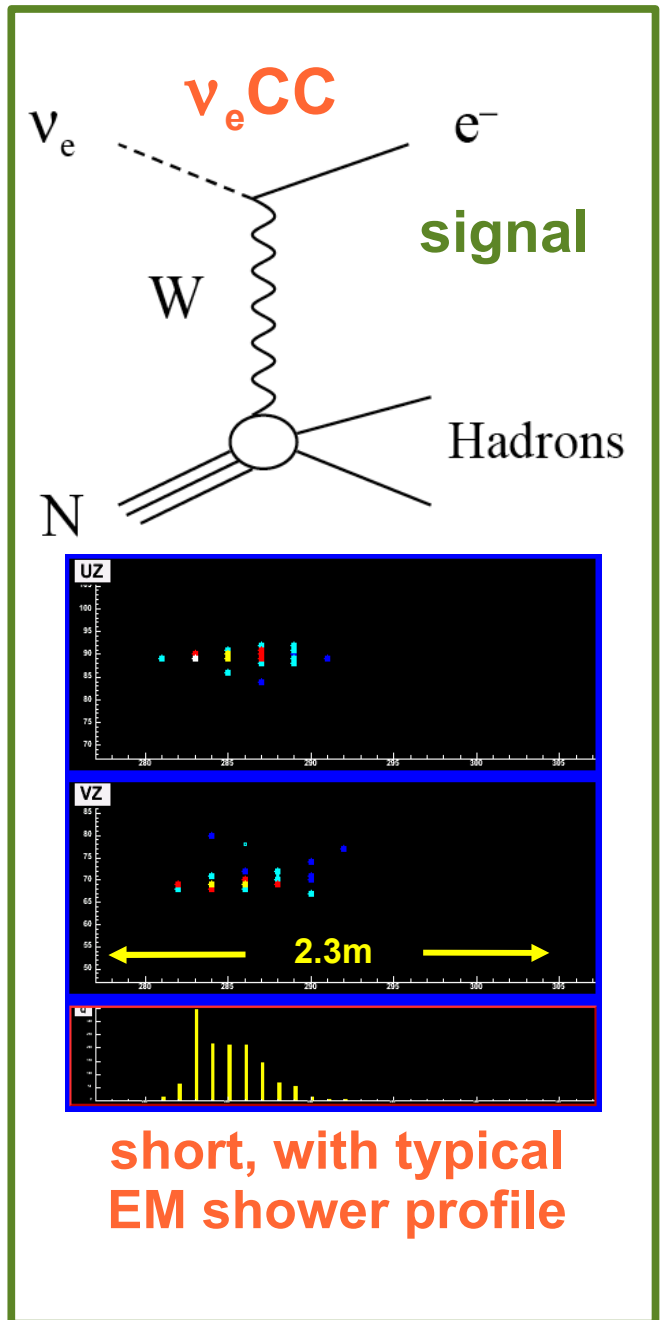
(MC event displays)



long  $\mu$  track+  
hadronic activity at  
vertex



short event, often  
diffuse

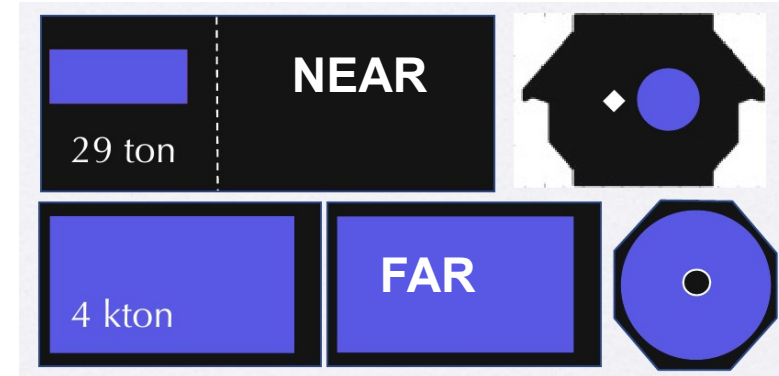


short, with typical  
EM shower profile

# Event Selection

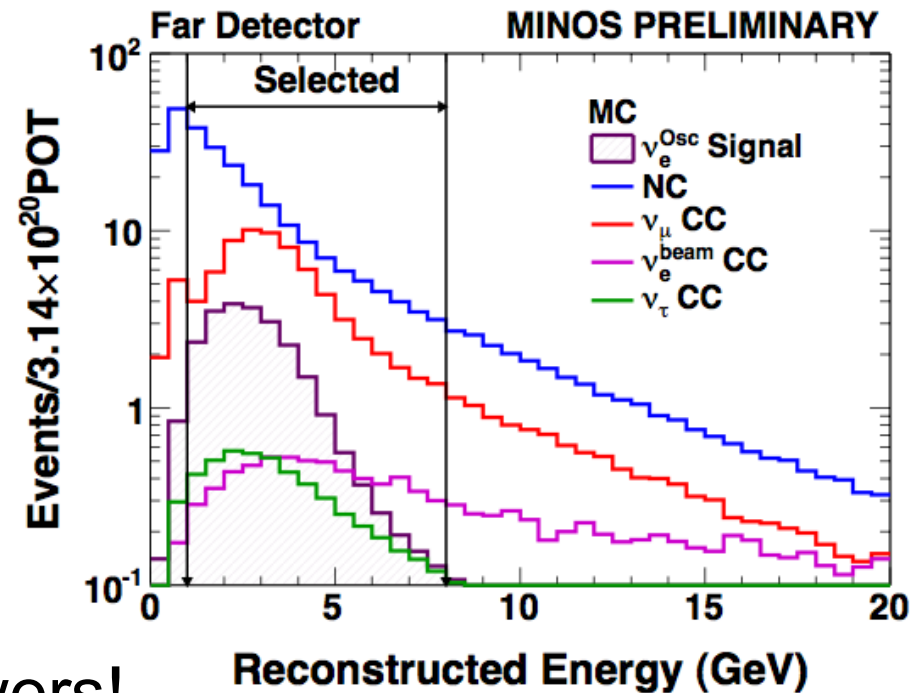
## Select good beam-induced events:

- Detector quality and beam quality
  - Fiducial volume
  - Reject cosmic tracks
- (selects 1188 far detector data events)



## Preselection: reduce background

- reject events with a long track
  - at least 1 shower
  - at least 4 hit planes in a row
  - reco energy 1-8 GeV
  - improves signal:background from 1:55 to 1:12
- (selects 227 far detector data events)



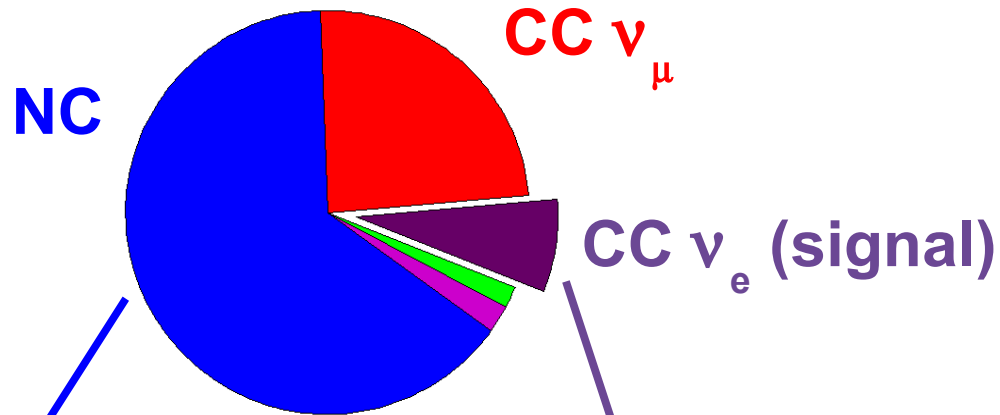
**$\nu_e$  selector:** select the  $\nu_e$ -like showers!



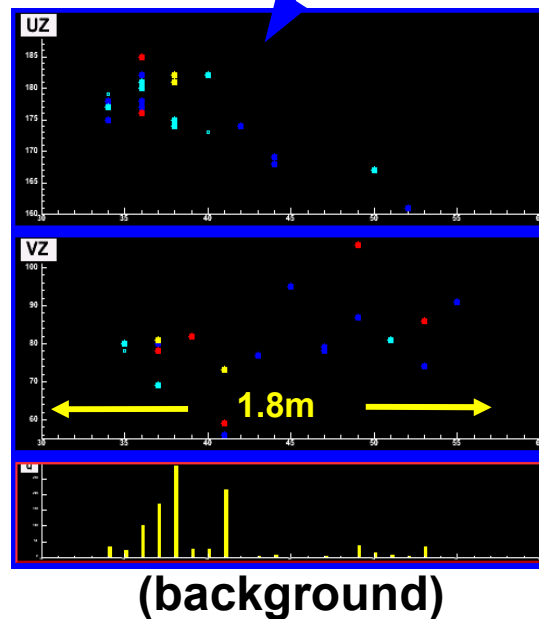
# Selecting $\nu_e$ -like showers

After preselection, the background is mostly NC

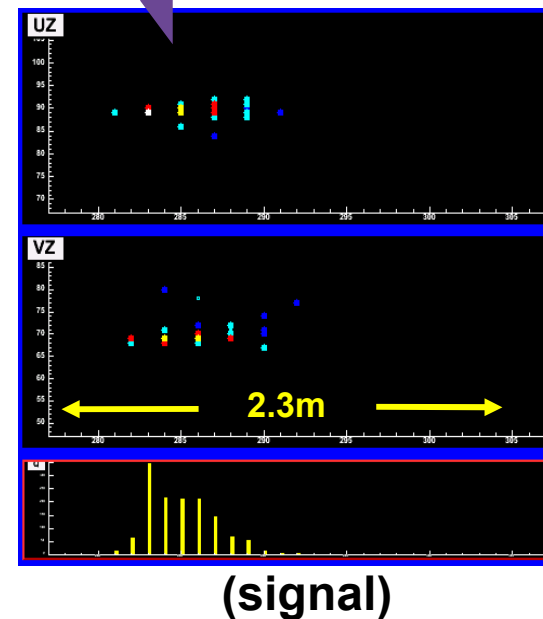
Far detector  
breakdown after  
preselection:



Need a  
method to  
discriminate



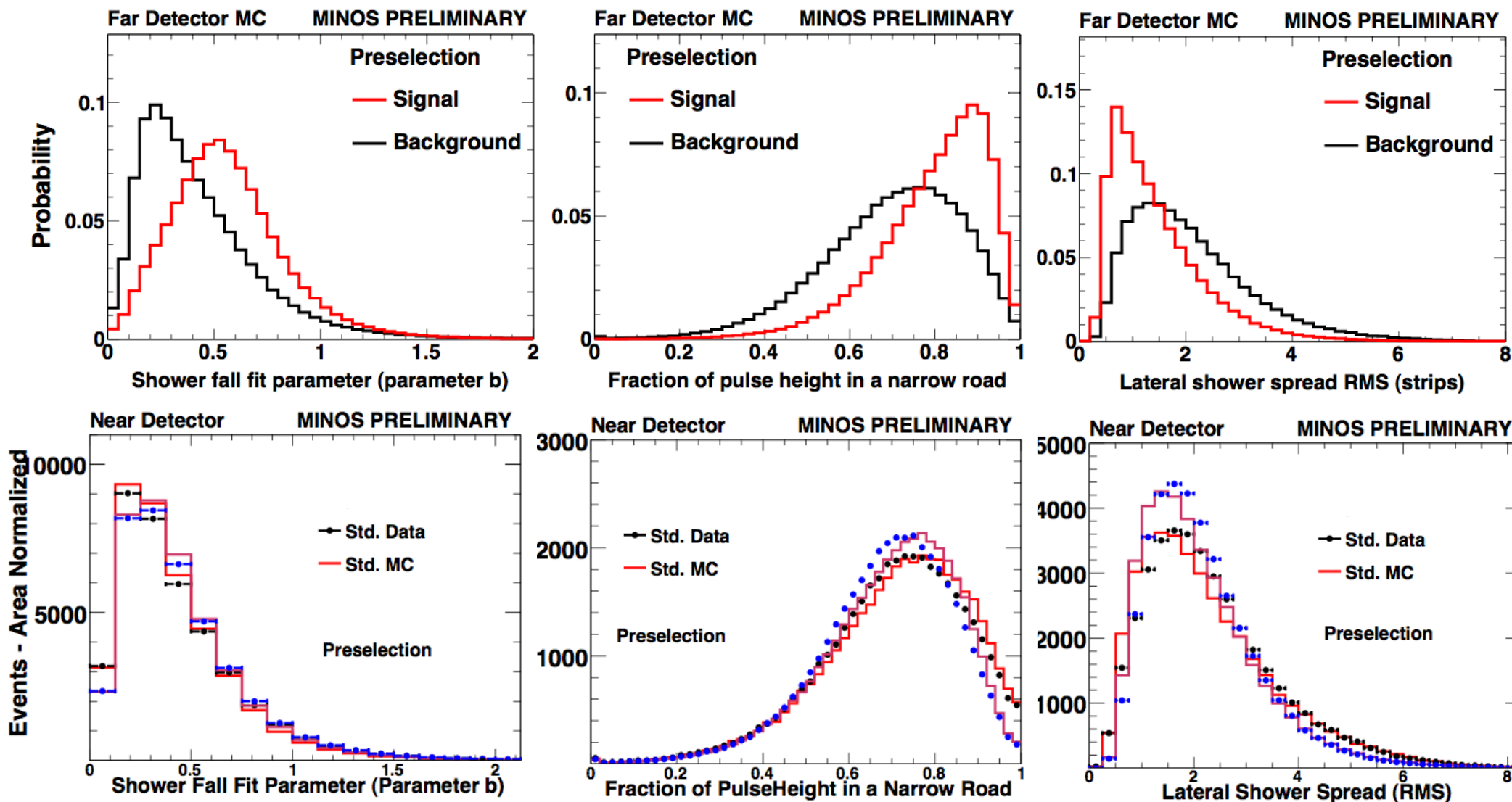
and



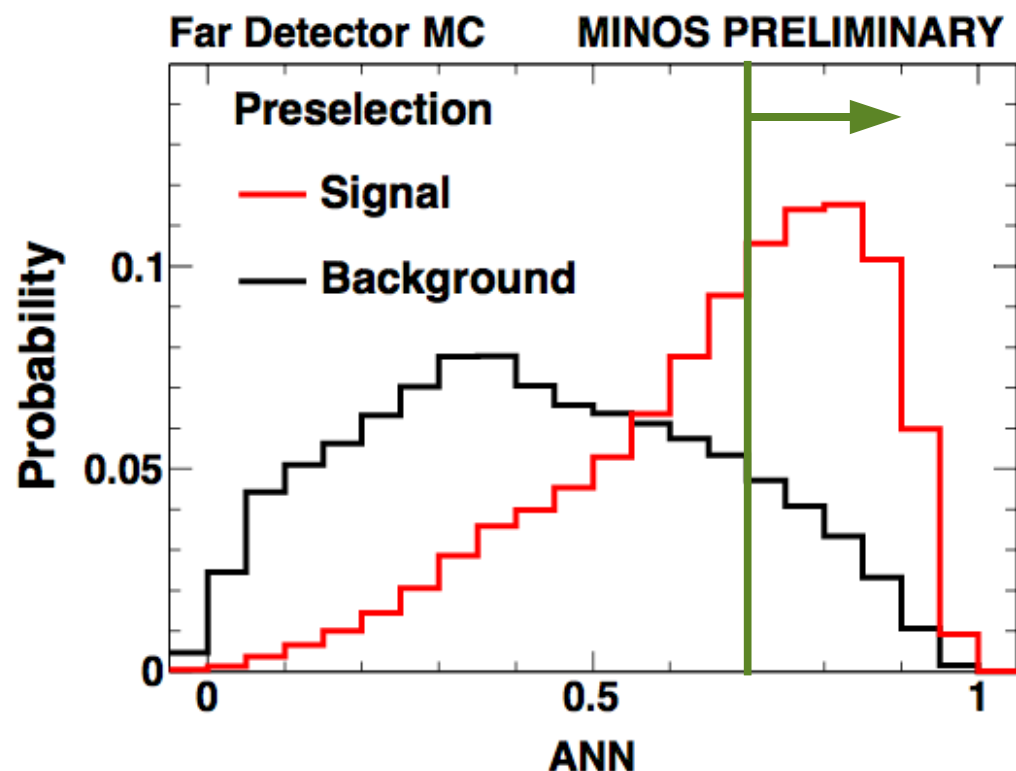
# Artificial Neural Net (ANN)

11 input variables that characterize the shower shape

some of the best variables....



# Artificial Neural Net (ANN)



**With a cut of  $ANN > 0.7$ :**

signal efficiency 41%

NC rejection >92.3%

CC rejection >99.4%

signal/background 1:4

this is our primary  $\nu_e$  selection method

# Library Event Matching (LEM)

$\nu_e$  selection by hit pattern recognition

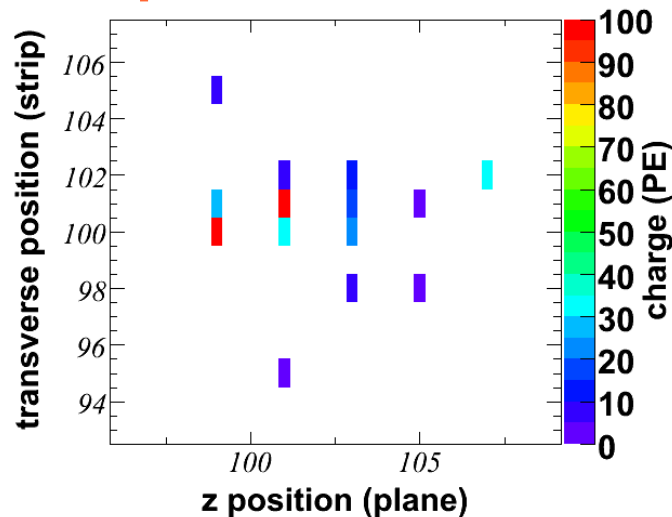
an alternative to ANN

Create a library of CC  $\nu_e$  and NC events

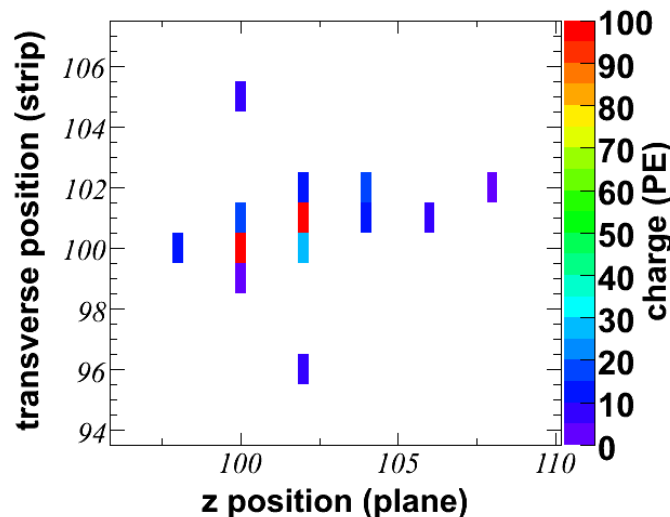
For each input event, select the 50 best matches from the library and construct a likelihood based on:

- the fraction of 50 best matches that are CC  $\nu_e$
- mean  $y$  of the CC  $\nu_e$  best matches ( $y$  = fraction of energy given to hadrons)
- mean fraction of charge in shared strips of the CC  $\nu_e$  best matches

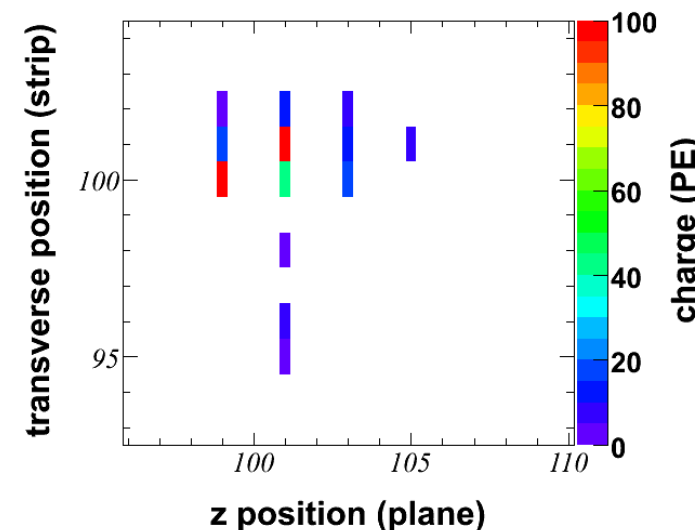
input data event



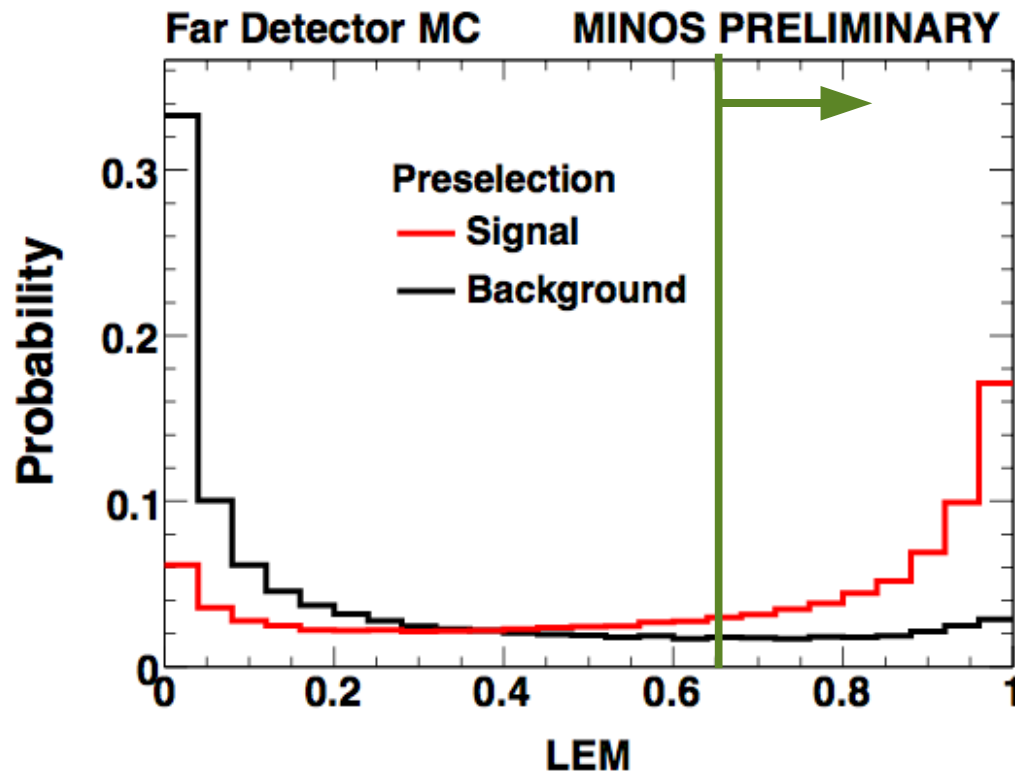
good match



bad match



# Library Event Matching (LEM)



**With a cut of  
LEM > 0.65:**

signal efficiency 46%  
NC rejection >92.9%  
CC rejection >99.3%  
signal/background 1:3

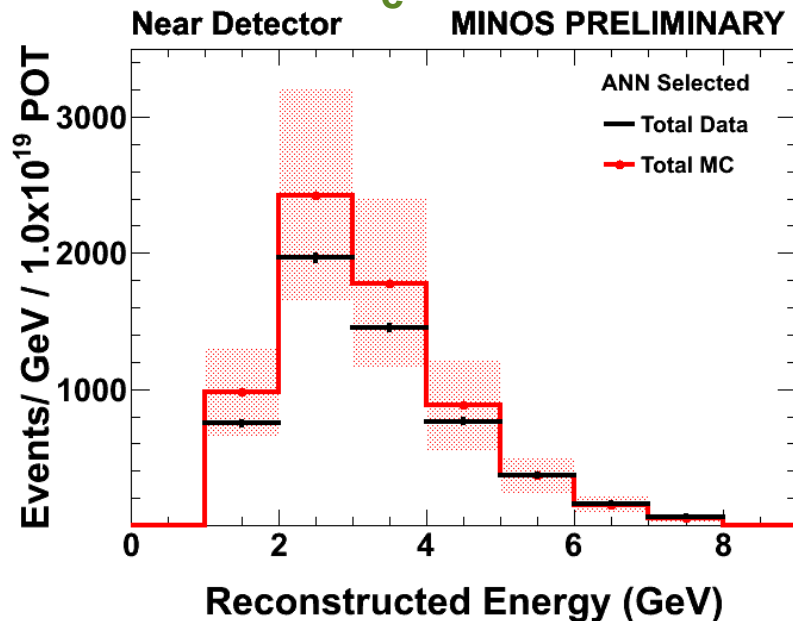
this is our secondary  $\nu_e$  selection method



# ND data to FD prediction?

Now that we have the  $\nu_e$  selection criteria, how do we make the far detector prediction?

## ANN-selected near detector $\nu_e$ candidates



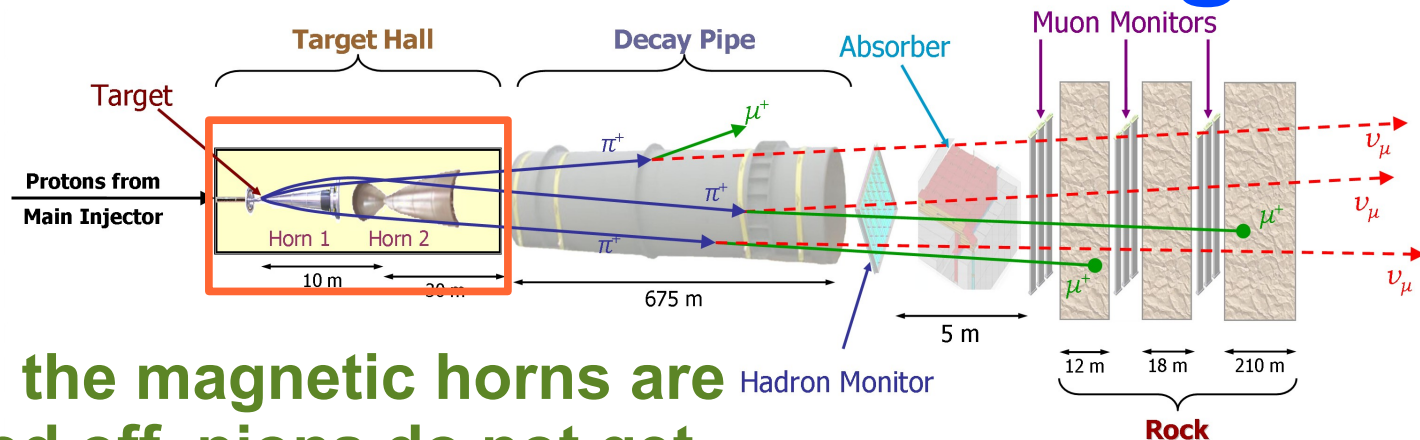
To first order, need only account for differences in **flux** ( $\sim 1/R^2$ ) and **fiducial volume**.

For more accurate extrapolation, need to consider **oscillation** ( $\nu_\mu$  disappearance affects CC component), **detector effects**, etc.

**Need to separate the near detector data into CC  $\nu_\mu$ , NC, beam  $\nu_e$  components.**

We can use horn-off data to do this...

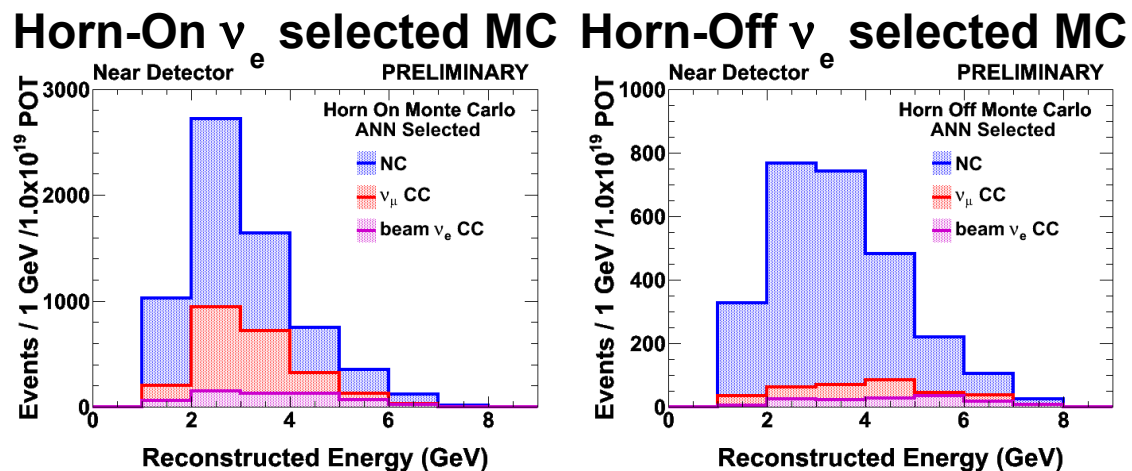
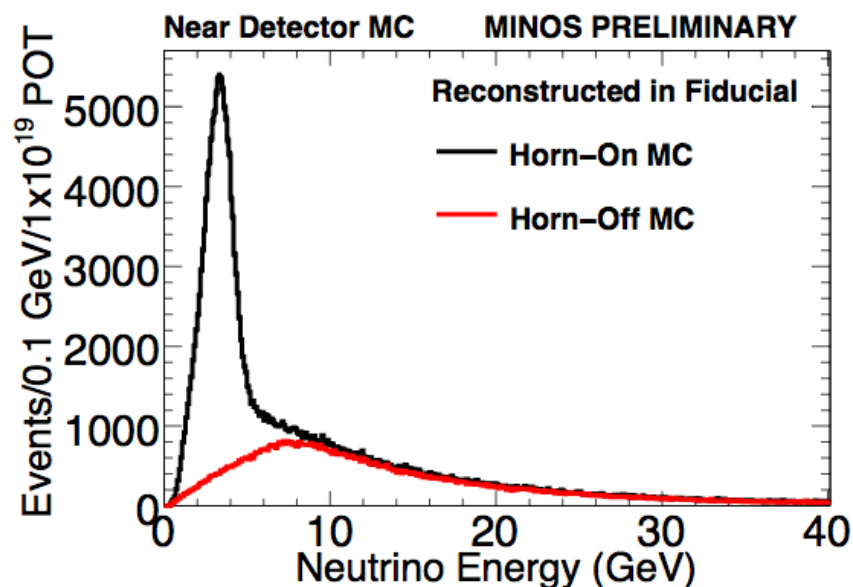
# Horn-Off Beam Configuration



5.5e18 POT  
of data in this  
configuration

When the magnetic horns are turned off, pions do not get focused, and the low energy peak of the neutrino energy spectrum disappears.

There is less contamination from low energy (short track) CC  $\nu_\mu$  events in the  $\nu_e$  selected near detector data.



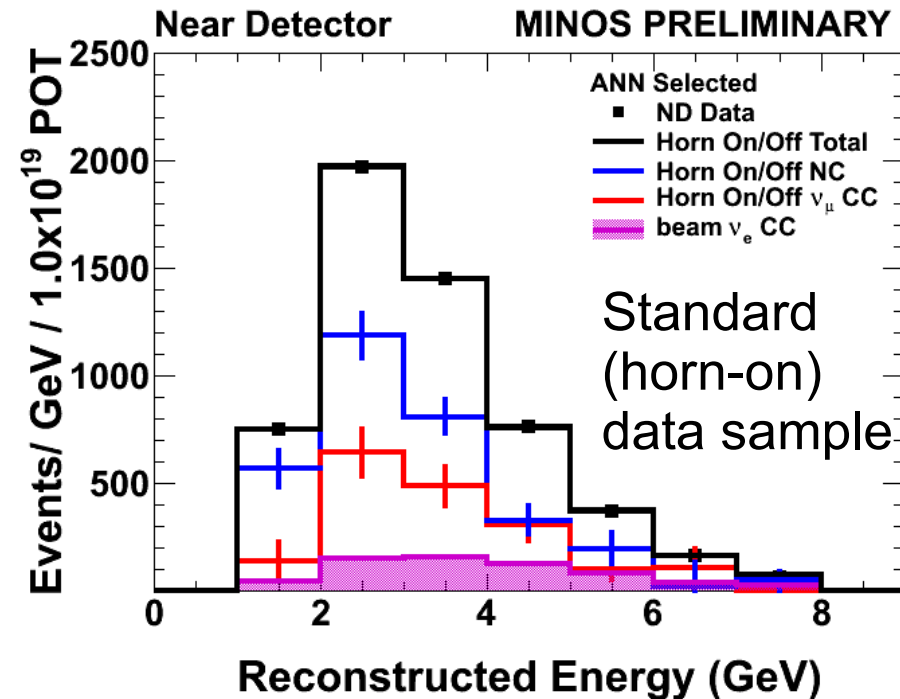
# Data-Driven Background Separation: Horn-On/Horn-Off

How does horn-off data help us?

The horn-off to horn-on ratio of selected CC and NC events is well-modeled in the MC

$$r_{NC} = N_{NC}^{OFF} / N_{NC}^{ON}$$

$$r_{CC} = N_{CC}^{OFF} / N_{CC}^{ON}$$



We have two data samples with different (and unknown) CC and NC components. But we do know (from the MC) the *relative* number of CC and NC events between the two samples.

→ we can calculate the individual components using the ratios

(The beam  $\nu_e$  component is taken from MC)

# Muon Removal Technique

ND Data Event

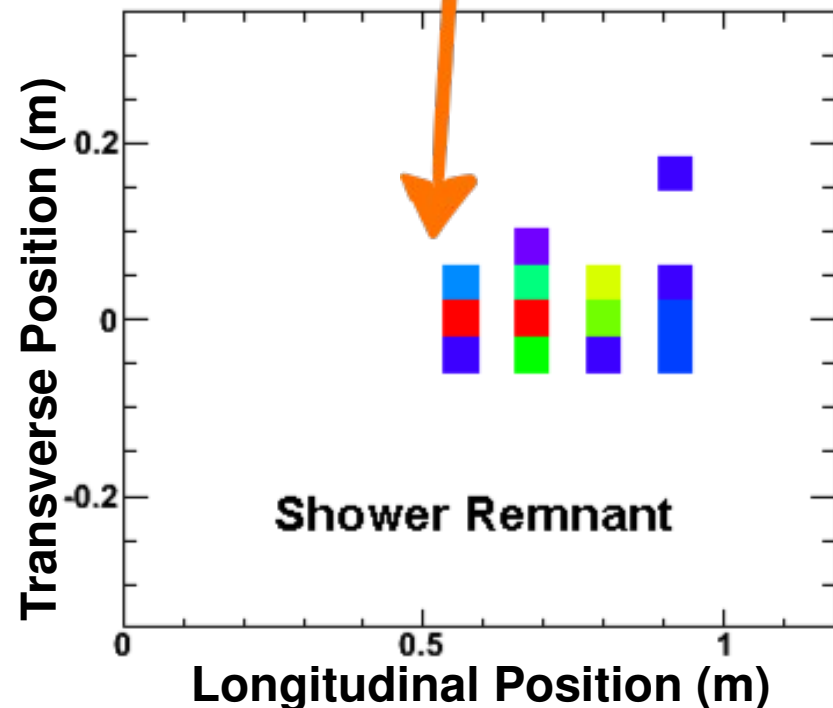
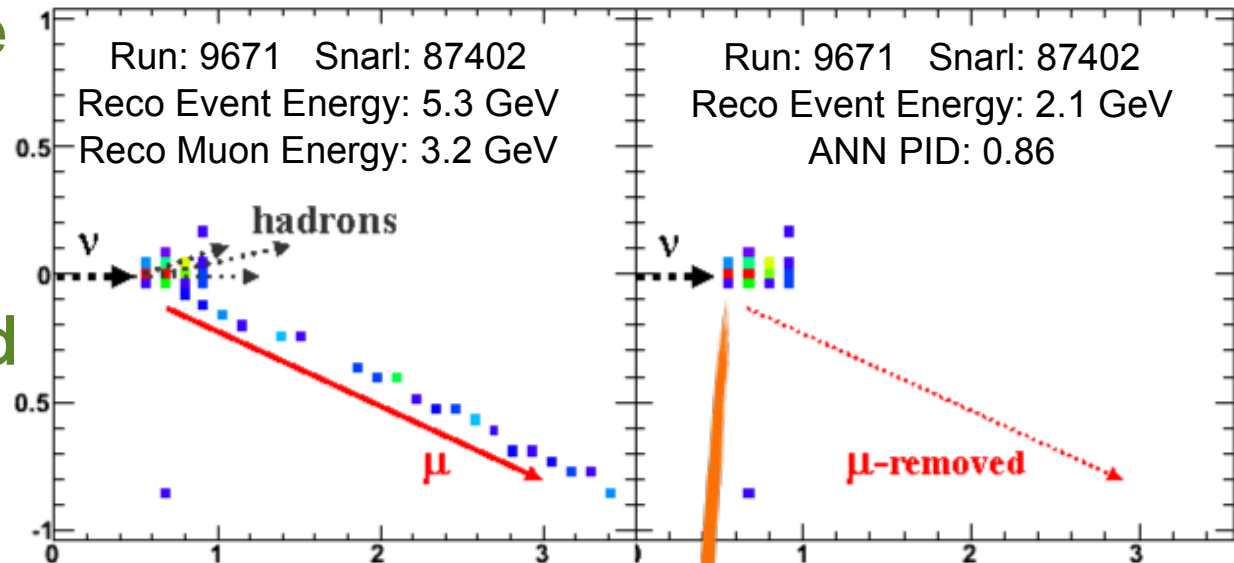
As a cross-check for the  
Horn on/off background  
analysis...

## Muon Removed Charged Current (MRCC)

- 1) take good CC events
- 2) remove the hits  
associated with the muon
- 3) re-reconstruct the  
hadronic shower

can do this for data and MC

an independent sample to  
study hadronic showers



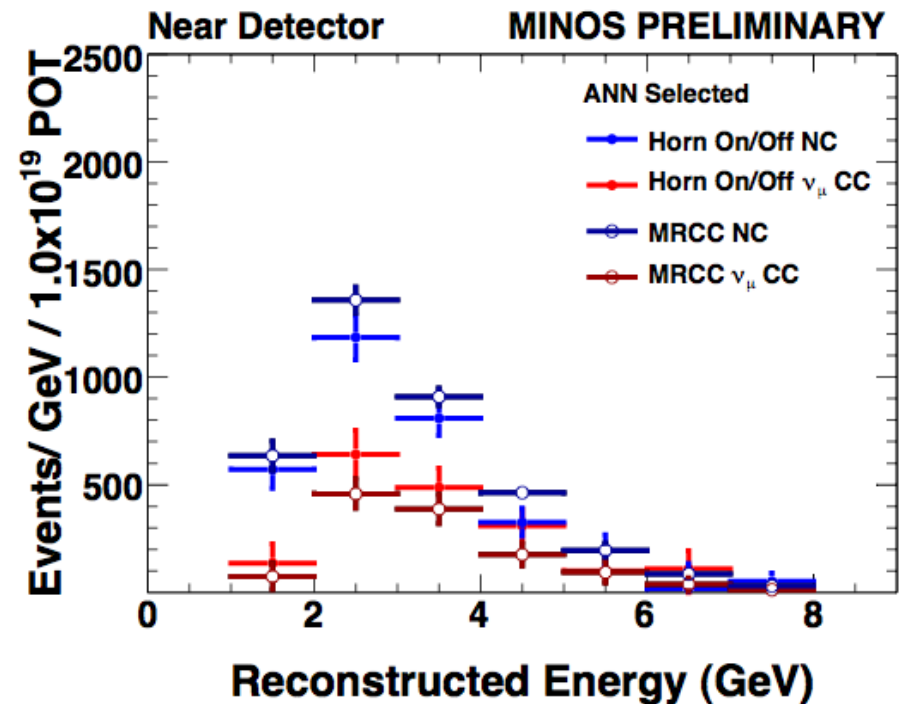
# MRCC

As a cross-check for the  
Horn on/off background  
analysis...

Apply  $\nu_e$  selection to the  
MRCC data and MRCC MC  
and use data/MC ratio to  
predict NC component of  
standard data selection

$$NC_i^{corr} = \frac{MRCC_i^{data}}{MRCC_i^{MC}} \times NC_i^{MC}$$

Results are consistent with  
Horn on/off method





# Extrapolation to Far Detector

## Far/Near Ratio:

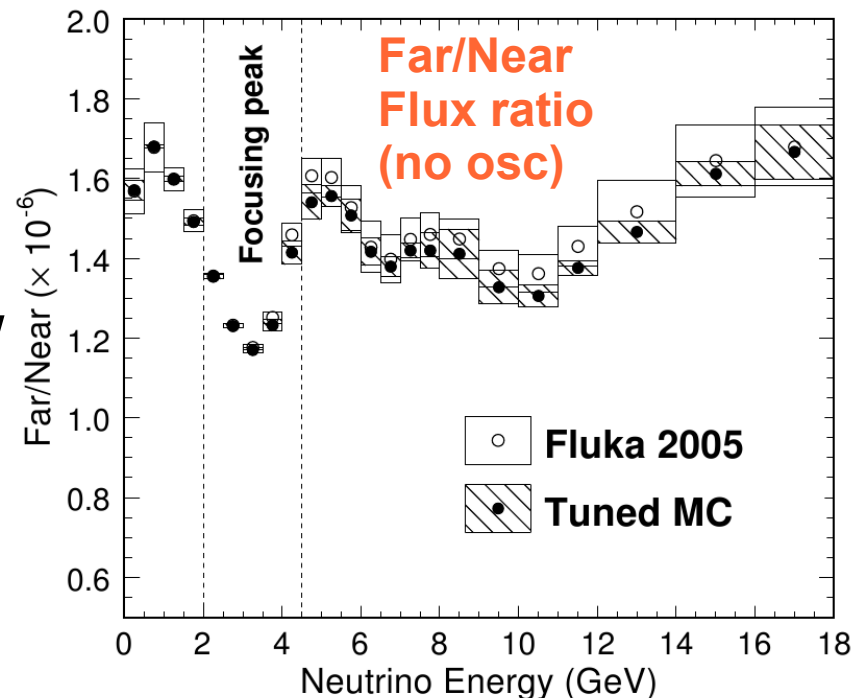
Ratio of MC  $\nu_e$  selected events in bins of reco energy

far detector background prediction  $F_i^{predicted, \alpha} = N_i^{\alpha} \times \left( \frac{f_i^{\alpha}}{n_i^{\alpha}} \right)$  Far/Near ratio  
near detector selected data

**Far/Near ratio** accounts for:

Largest effects

- • Flux ( $1/R^2$ , geometry, focusing, acceptance, decay kinematics)
- • Fiducial volume (4000 tons/29 tons)
- energy smearing
- $\nu_{\mu}$  disappearance
- detector effects (next slide)
- etc



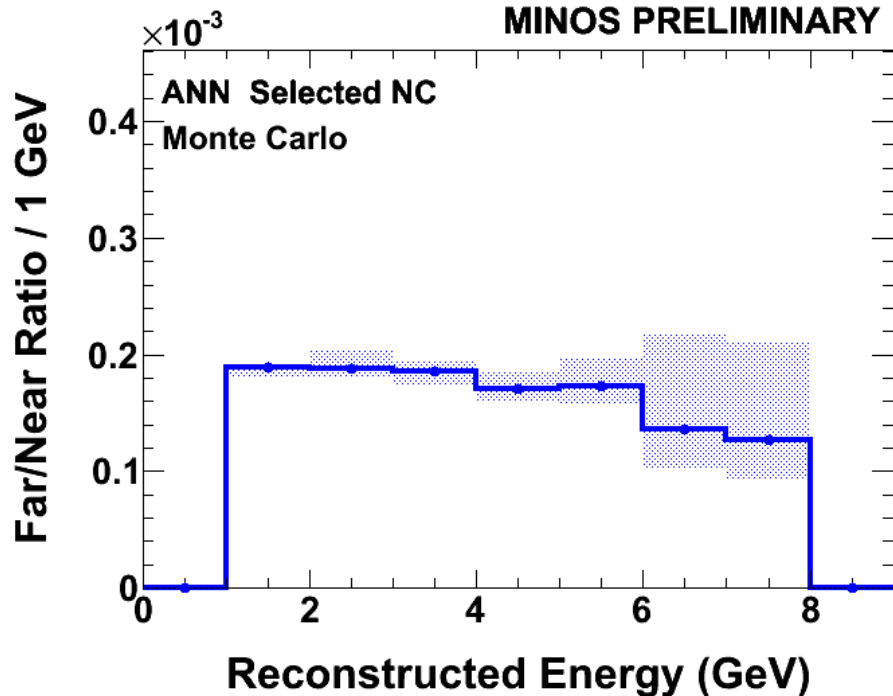
# Far/Near Differences

**Far/Near differences that are taken into account by ratio:**

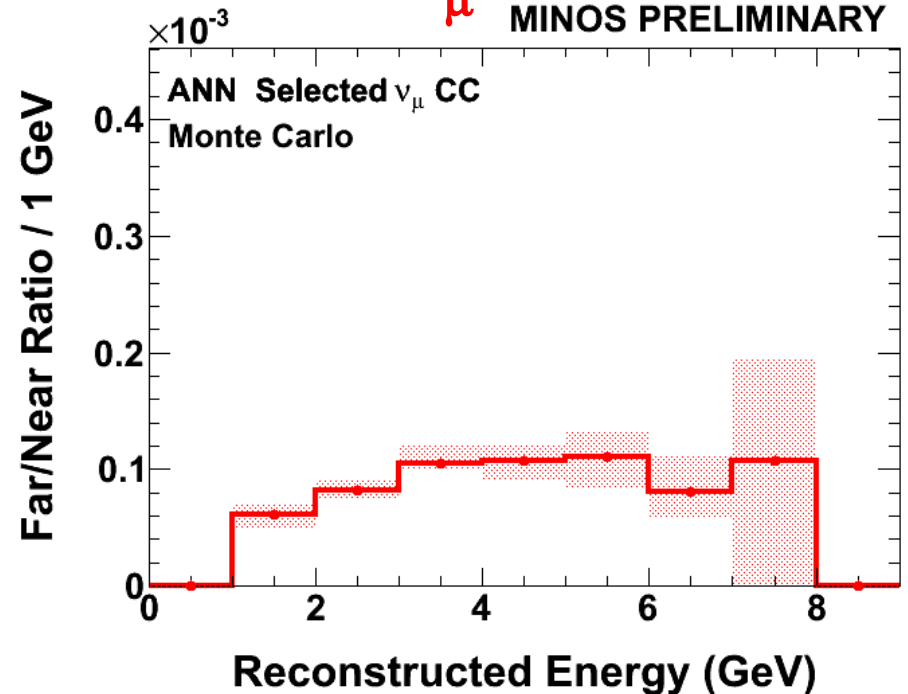
- ♦ difference in fiber length (light level difference)
- ♦ multiplexing in the far detector (8 fibers per PMT channel)
- ♦ one-sided readout in the near detector
- ♦ PMTs (64-channel in near, 16-channel in far) - different crosstalk pattern, gains, front end electronics
- ♦ faster readout in near detector
- ♦ relative energy calibration

# Far/Near Ratio

NC



CC  $\nu_\mu$



Far detector beam  $\nu_e$  prediction is taken from MC.

Predictions for oscillated CC  $\nu_\tau$  and CC  $\nu_e$  (signal) are made based on the CC  $\nu_\mu$  spectrum at the near detector.

# Systematic Uncertainties in Far Detector Background Prediction

**Extrapolation:**

**6.4%**

**Horn-On/Horn-Off**

**Method:**

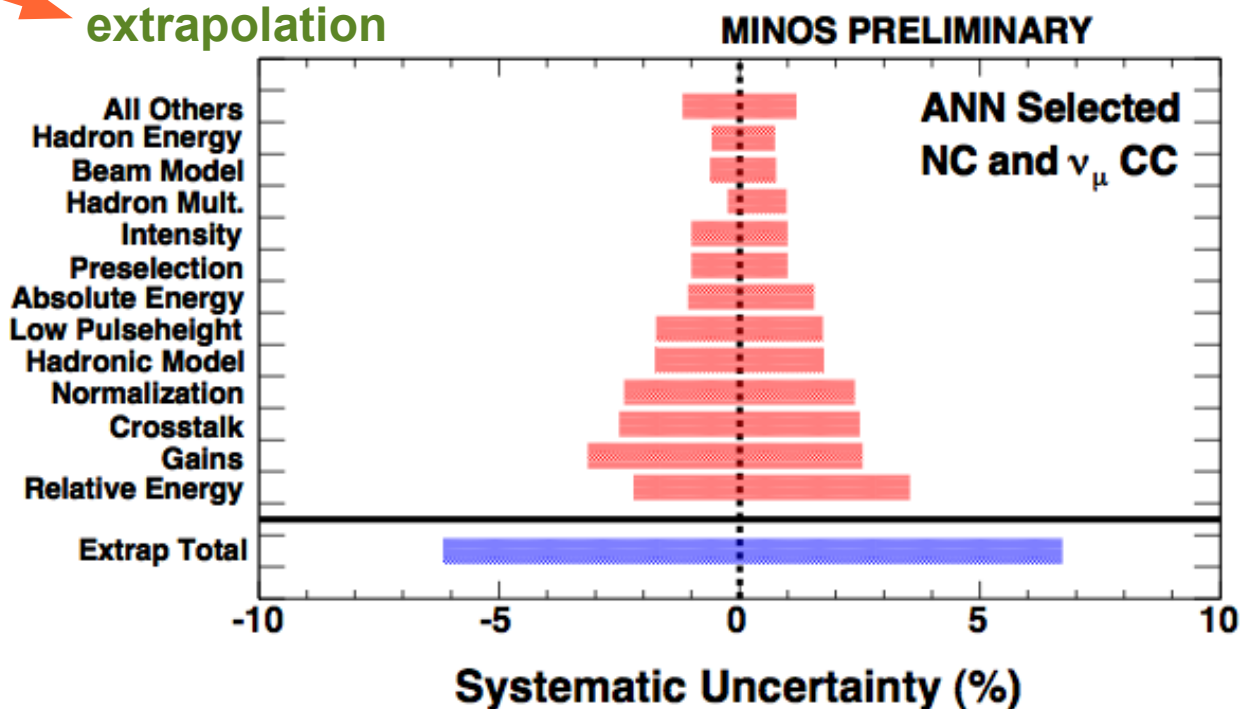
**3.5%**

**Total Systematic**

**Uncertainty:**

**7.3%**

Uncertainties related to extrapolation



# Far Detector Background Prediction

**Far Detector Background Prediction for  
3.14e20 POT:**

$$27 \pm 5(\text{stat}) \pm 2(\text{syst})$$

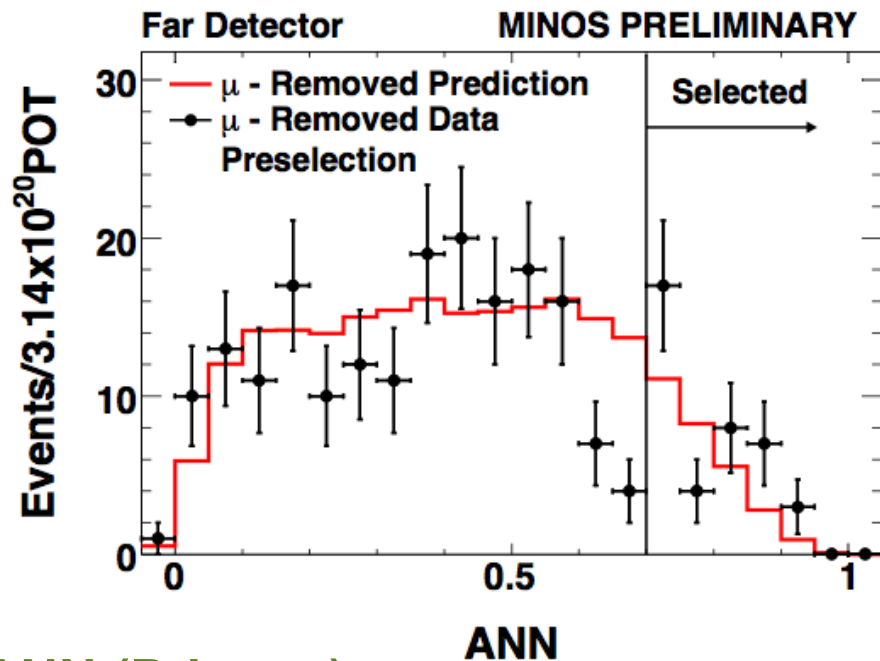
26.6 events: 18.2 NC, 5.1 CC  $\nu_\mu$ , 1.1 CC  $\nu_\tau$ , 2.2 beam  $\nu_e$

# Sidebands

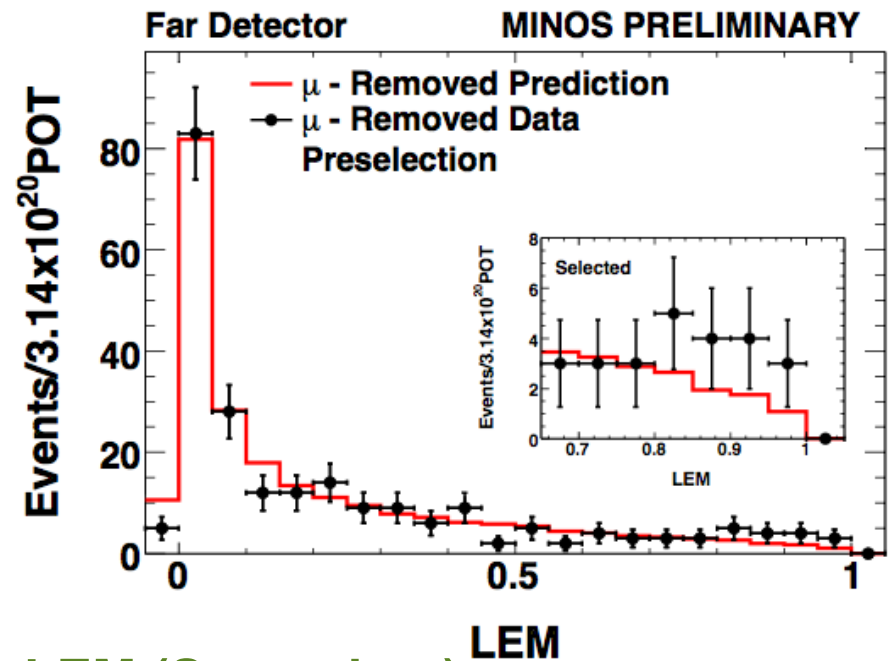
Before opening the box in the signal region, we examined three far detector data samples with no expected signal:

- far detector muon removed events
- far detector muon removed w/electron events
- far detector preselected events that fail the  $\nu_e$  selection cut

# Muon Removed Far Detector Events



**ANN (Primary):**  
 observe 39 events  
 expect  $29 \pm 5(\text{stat}) \pm 2(\text{syst})$

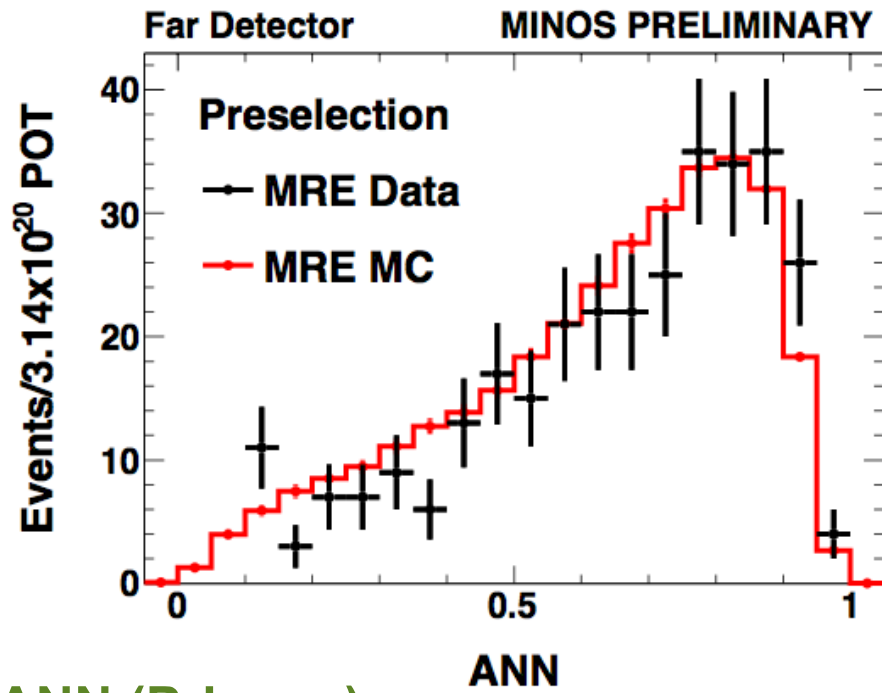


**LEM (Secondary):**  
 observe 25 events  
 expect  $17 \pm 4(\text{stat}) \pm 2(\text{syst})$

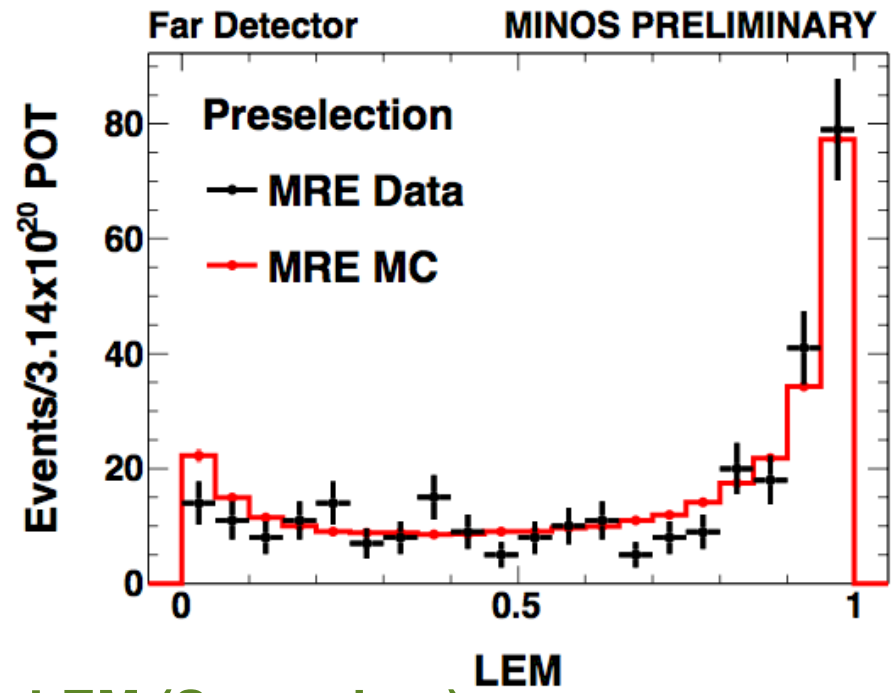
- Muon removed events were originally  $\nu_\mu$  CC events, so there is no signal in this sample.
- Good cross check that ANN/LEM behave as expected for hadronic showers (the major background)



# Muon Removed + Electron Far Detector Events



**ANN (Primary):**  
observe 159 events  
expect  $152 \pm 13(\text{stat}) \pm 12(\text{syst})$

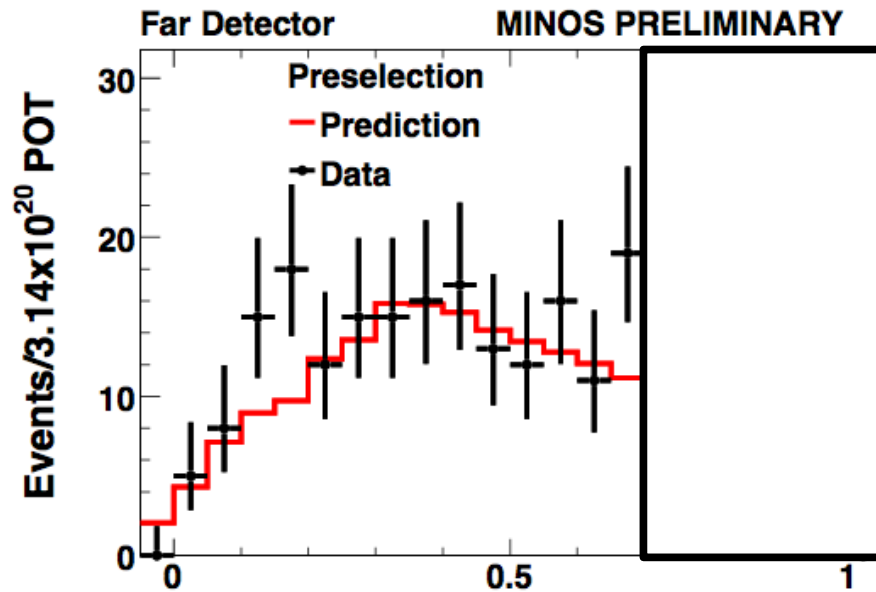


**LEM (Secondary):**  
observe 180 events  
expect  $176 \pm 13(\text{stat}) \pm 17(\text{syst})$

- Take muon removed events, add an electron and re-reconstruct
- Allows us to simulate signal with a real hadronic shower
- Good cross check that ANN/LEM behave as expected for signal-like events.

# Events Outside Signal Cut

ANN<0.7

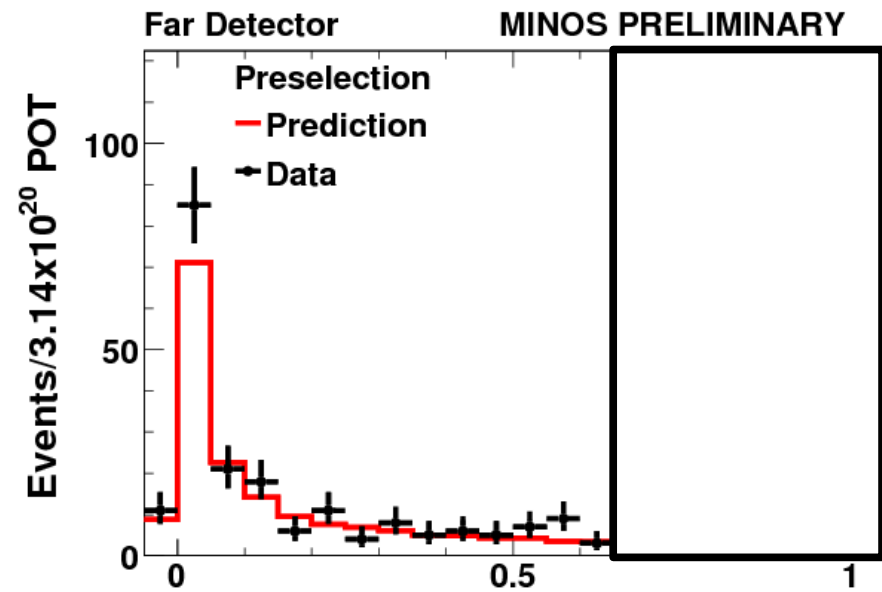


ANN (Primary):  
 $0 < \text{ANN} < 0.55$   
observe 146 events  
expect  $132 \pm 12(\text{stat}) \pm 8(\text{syst})$

$0.55 < \text{ANN} < 0.7$   
observe 46 events  
expect  $38 \pm 6(\text{stat}) \pm 2(\text{syst})$

- Good test of entire analysis chain - background prediction and extrapolation to far detector.

LEM<0.65



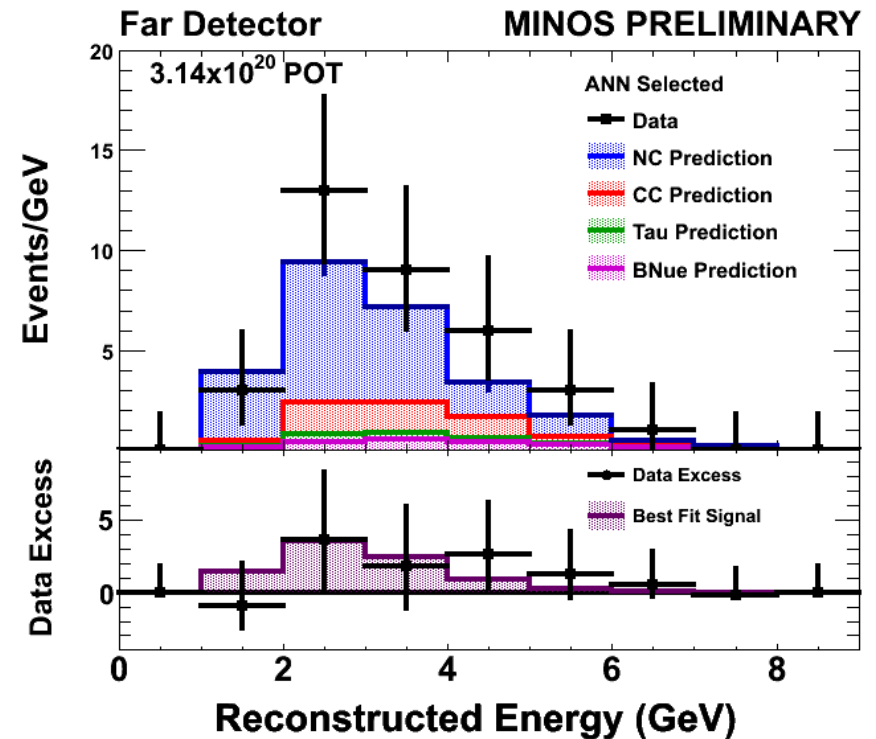
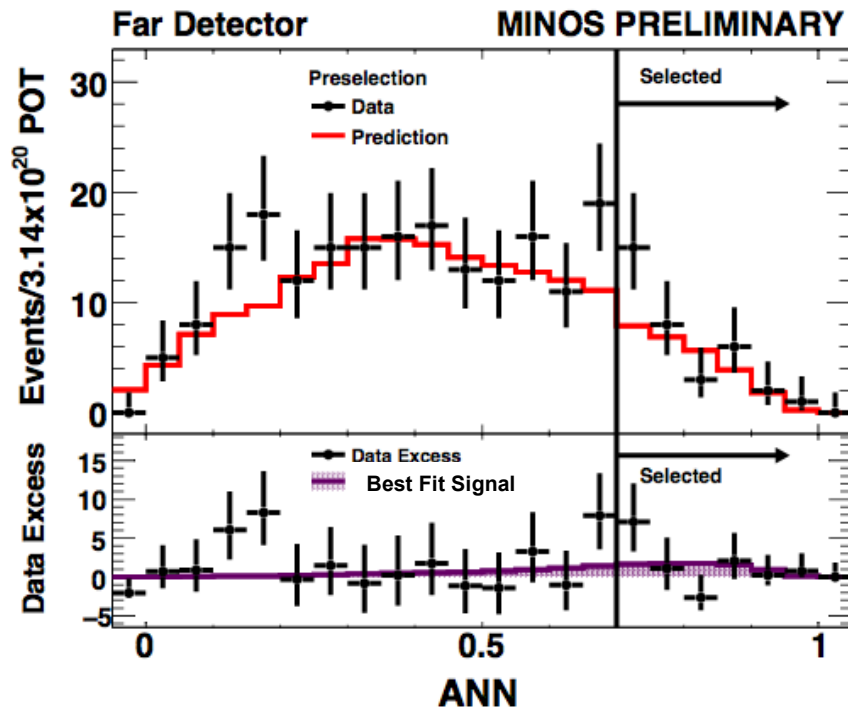
LEM (Secondary):  
 $0 < \text{LEM} < 0.55$   
observe 176 events  
expect  $157 \pm 13(\text{stat}) \pm 3(\text{syst})$

$0.55 < \text{LEM} < 0.65$   
observe 12 events  
expect  $7 \pm 3(\text{stat}) \pm 0(\text{syst})$

# $\nu_e$ Appearance Results for 3.14e20 POT

# $\nu_e$ Selected Far Data

## ANN (Primary Selection Method)

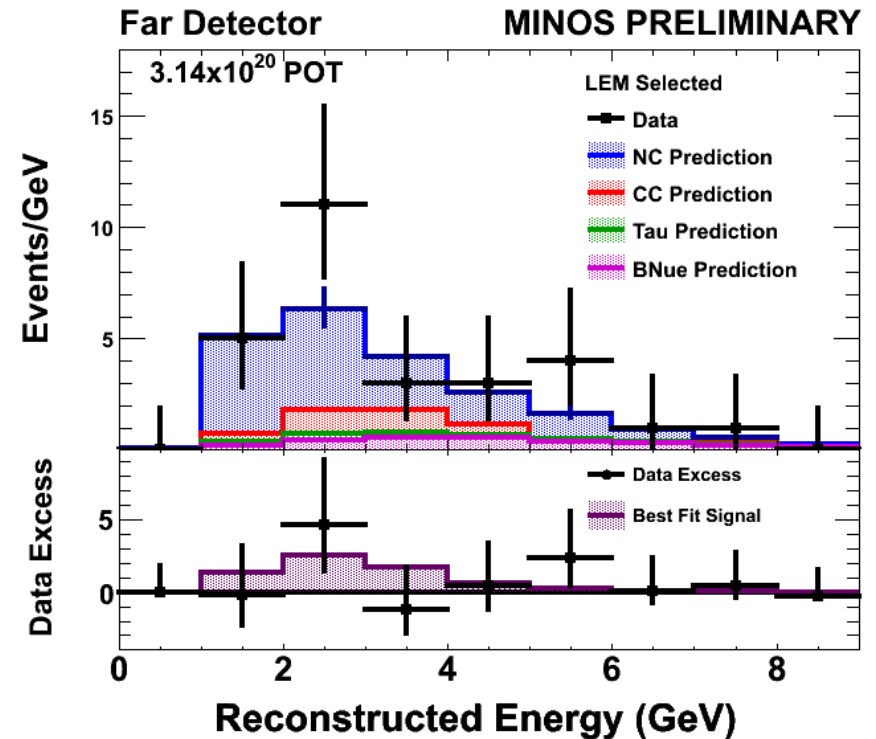
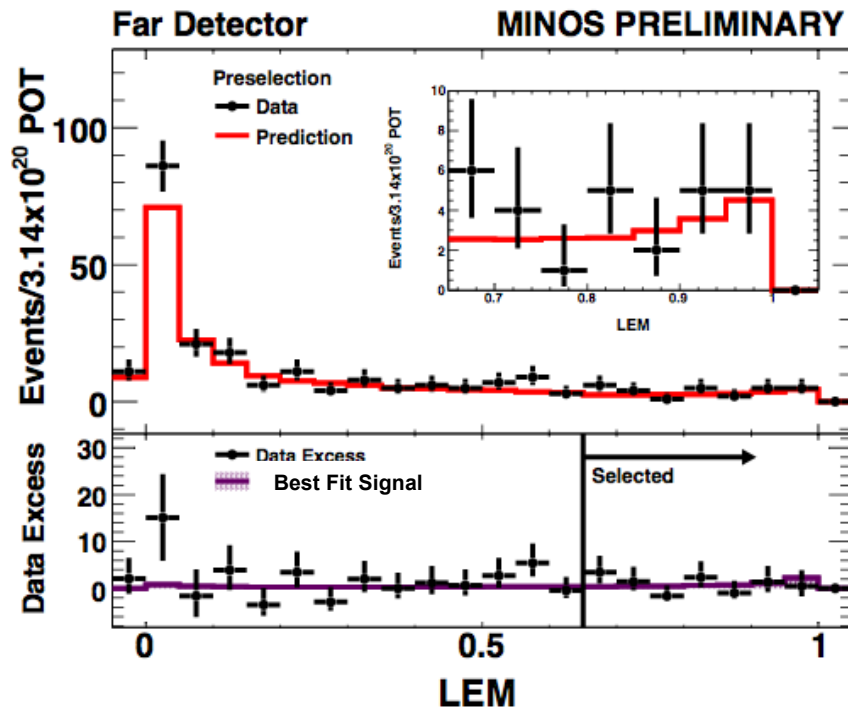


Observation: **35** events

Expected Background:  **$27 \pm 5(\text{stat}) \pm 2(\text{syst})$**  events

# $\nu_e$ Selected Far Data

## LEM (Secondary Selection Method)

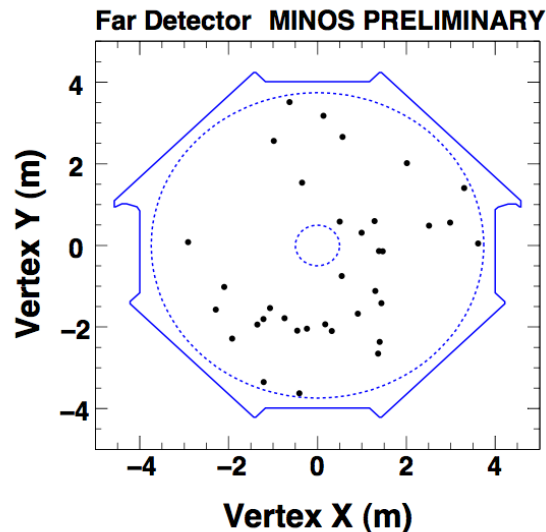


Observation: **28** events

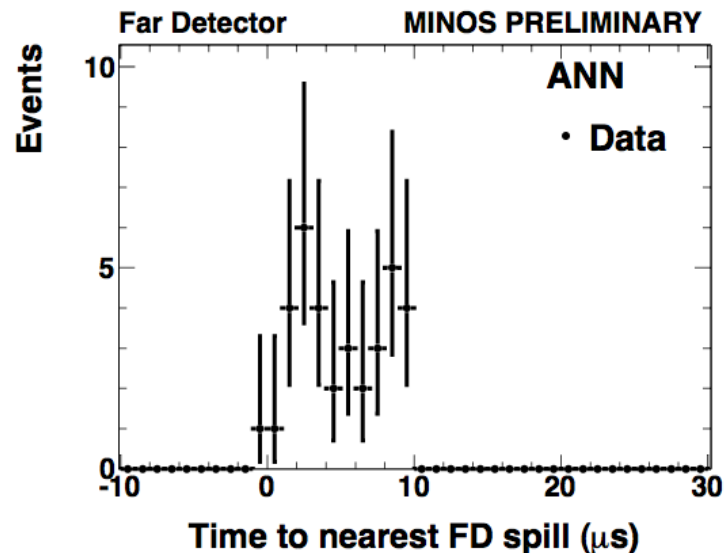
Expected Background: **22 ± 5(stat) ± 3(syst)** events

# Far Data Distributions

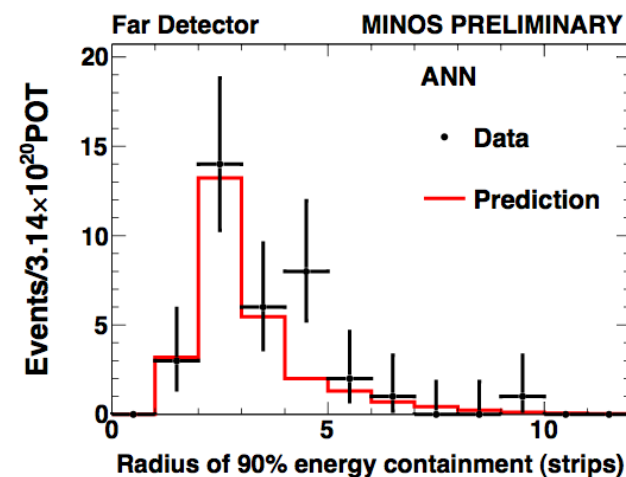
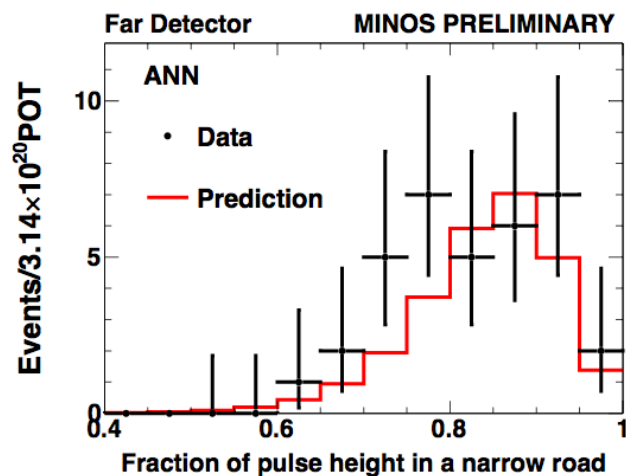
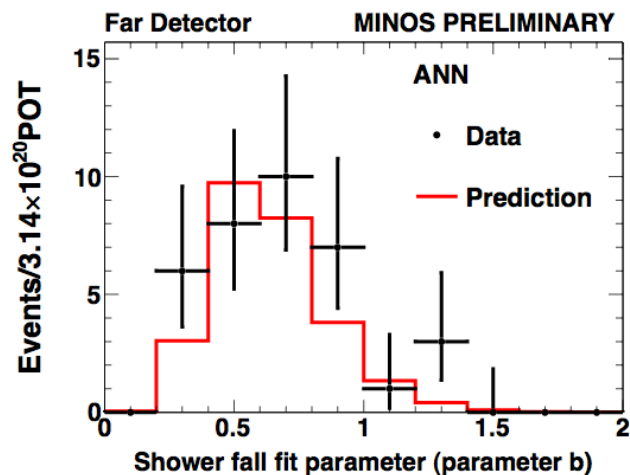
## Vertex



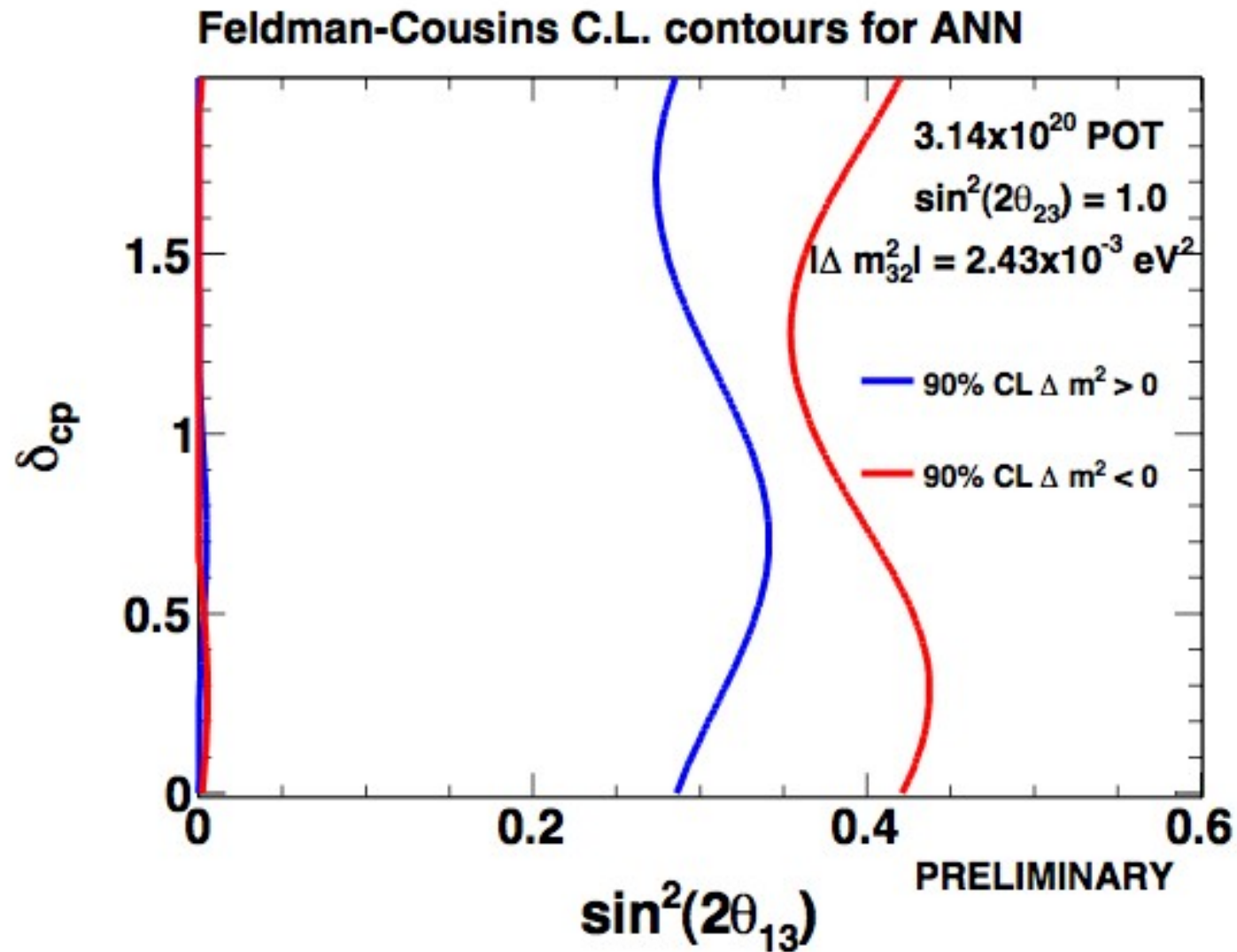
## Timing



## Some ANN variables

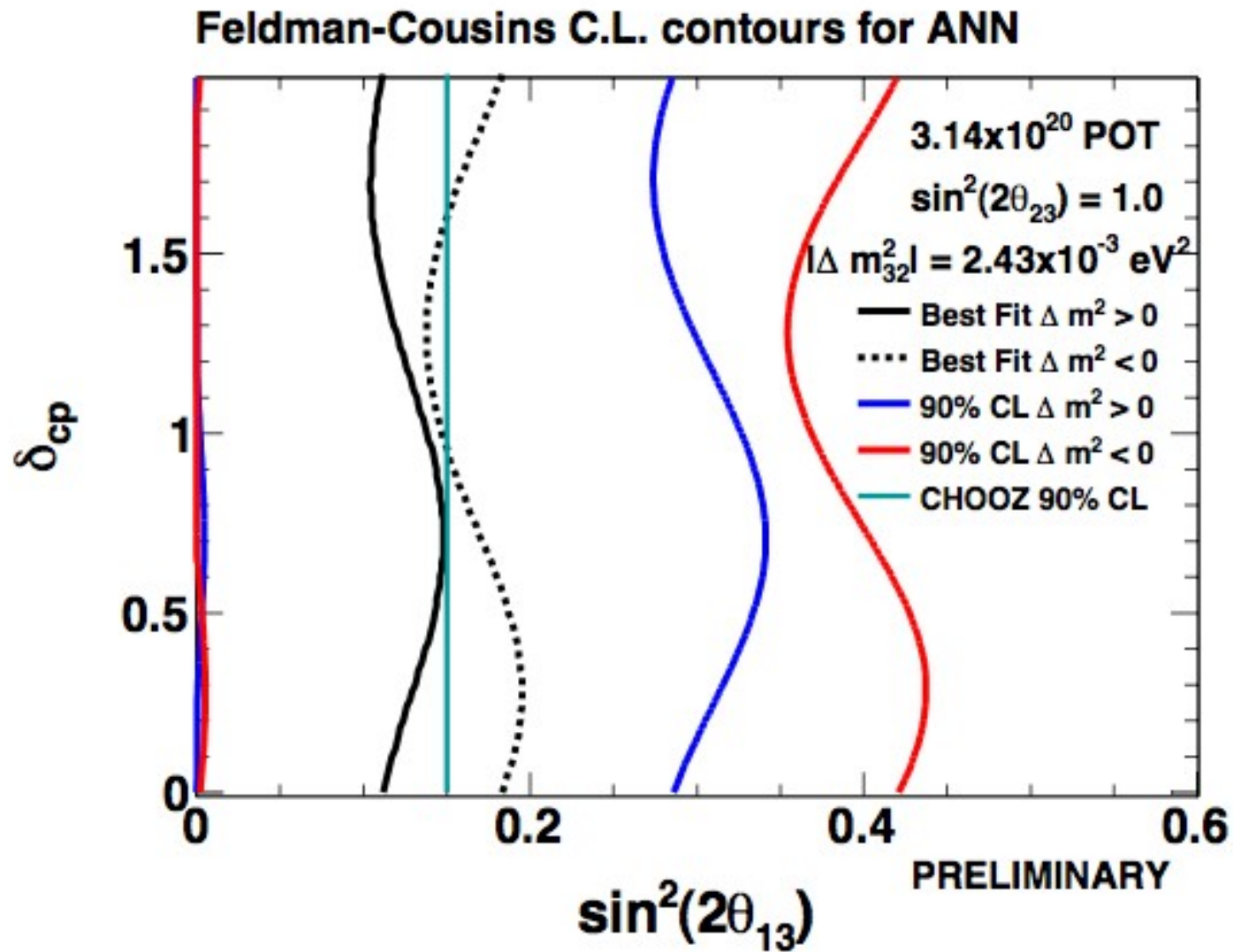


# 90% Confidence Level

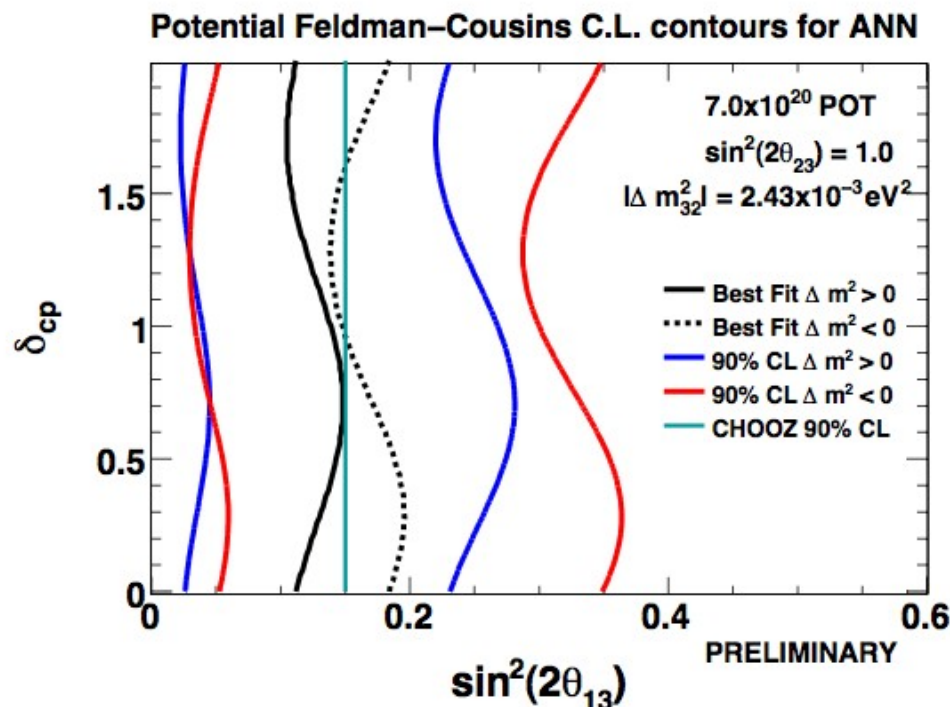




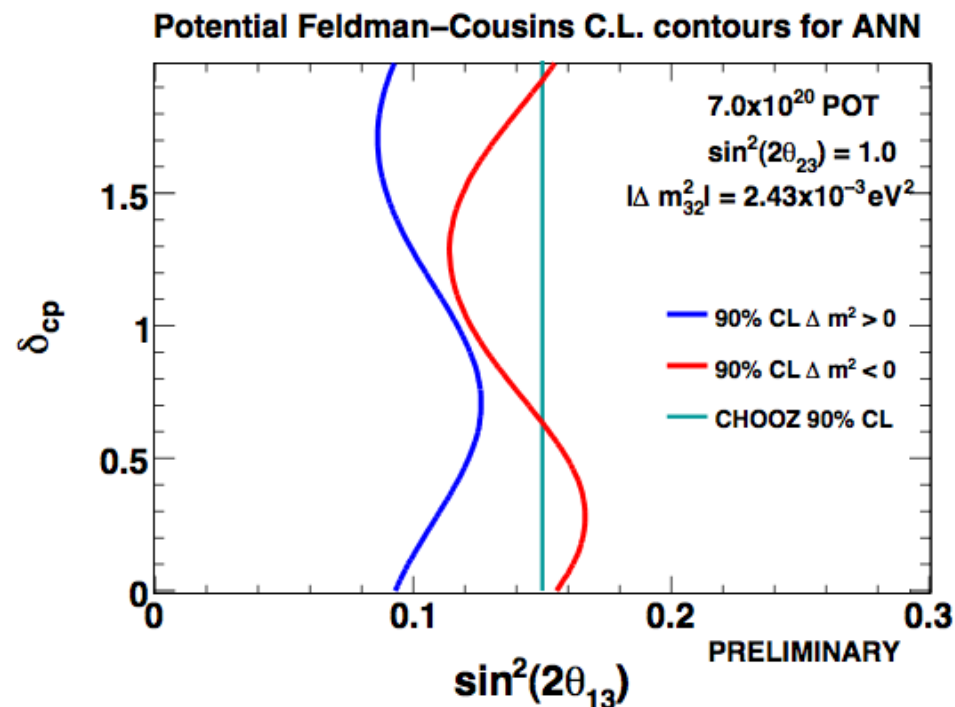
# 90% Confidence Level



# Future Prospects: 7e20 POT



Future result if the  
data excess  
persists



Future result if data  
excess goes away  
with more statistics

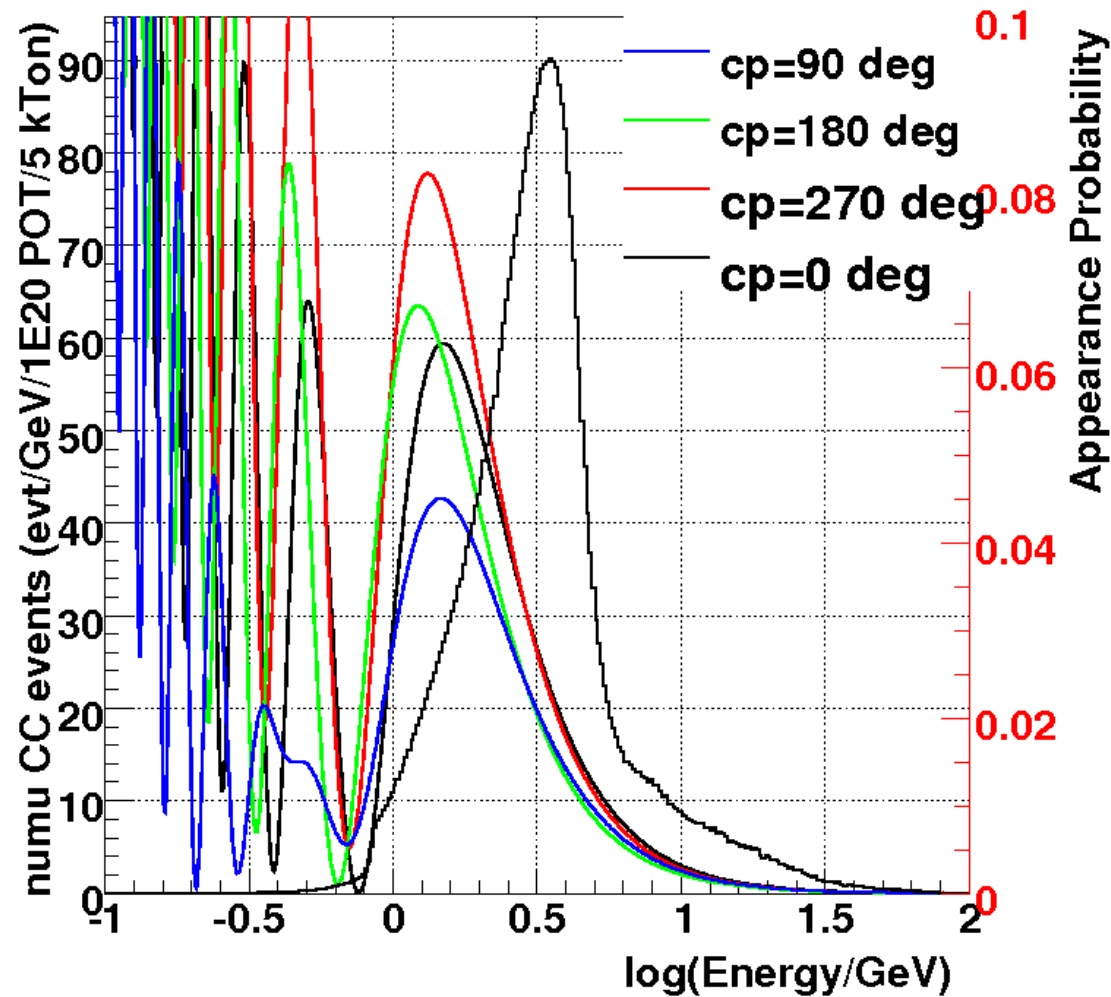
# Summary

- We have obtained our first results on the search for  $\nu_e$  appearance in MINOS
- We observed 35 events with a background expectation of  $27 \pm 5(\text{stat}) \pm 2(\text{syst})$  for  $3.14\text{e}20$  POT
- We set a 90% CL limit of  $\sin^2 2\theta_{13} < 0.29$   
(normal mass hierarchy,  $\delta=0$ )
- We are close to doubling this data in the current running – next results with  $>7\text{e}20$  POT!

# Backup Slides

# $\nu_e$ appearance at MINOS

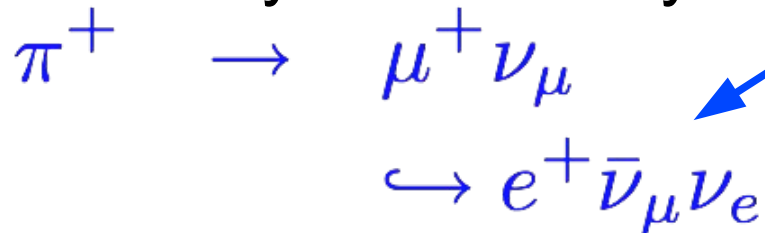
NuMI LE,  $\theta_{13}=0.16$  at 735km



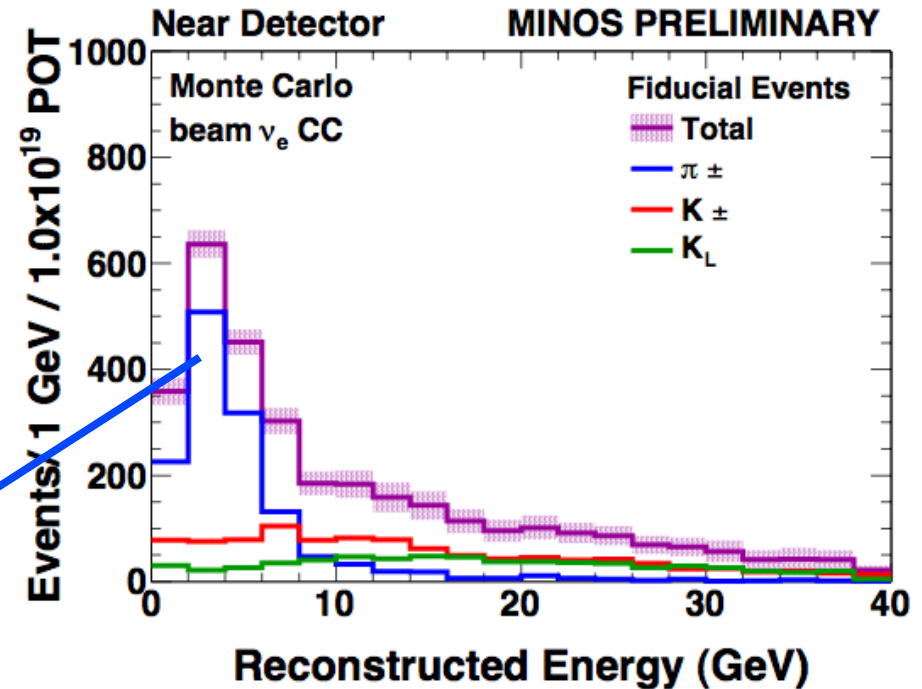
# Beam $\nu_e$ component

Neutrino beam has 1.3% of  $\nu_e$  contamination from pion and kaon decays.

Region of interest for the  $\nu_e$  oscillation analysis, 1-8 GeV, dominated by events from secondary muon decays:

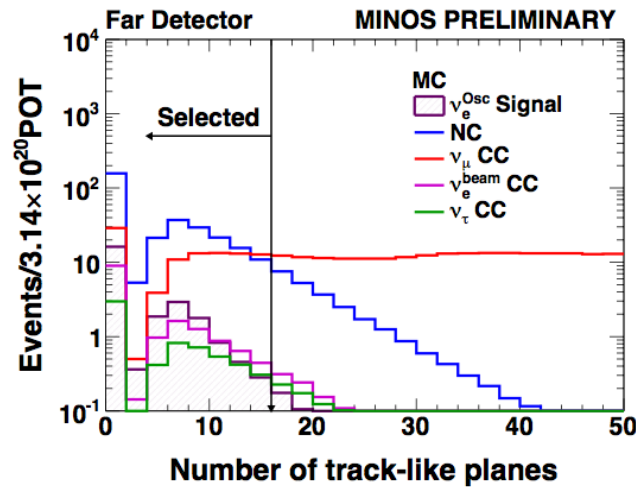
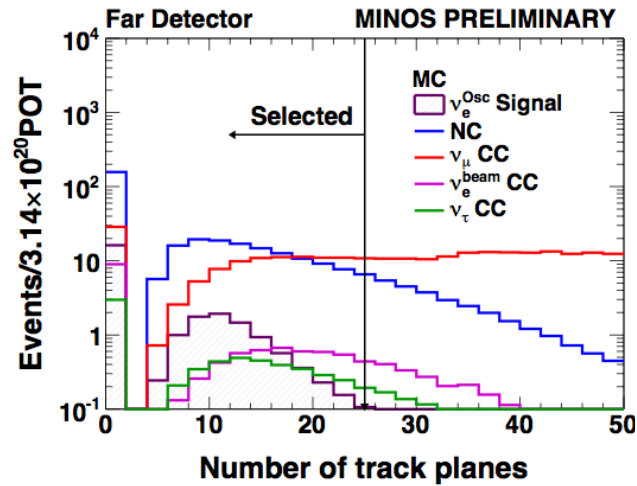


Near and Far beam  $\nu_e$  spectra are constrained by using  $\nu_\mu$  events from several beam configurations.

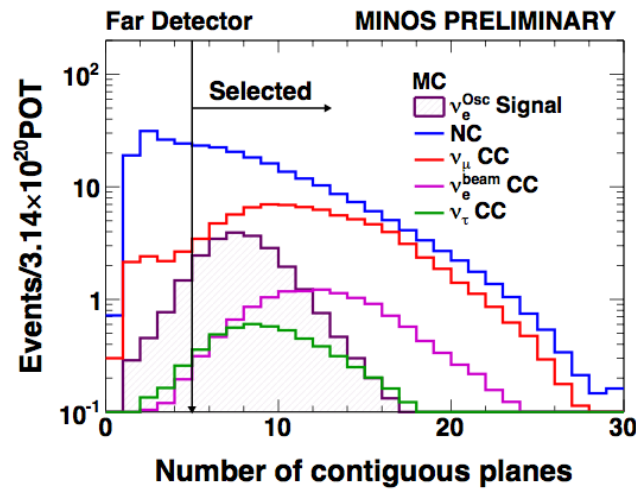
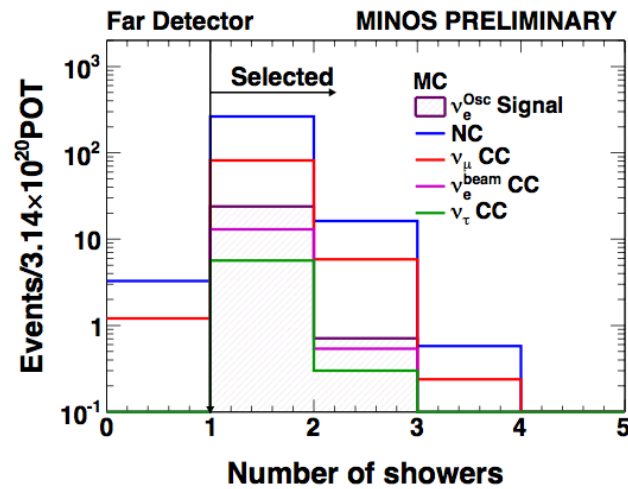


Uncertainties on the flux in the region of interest are ~10%

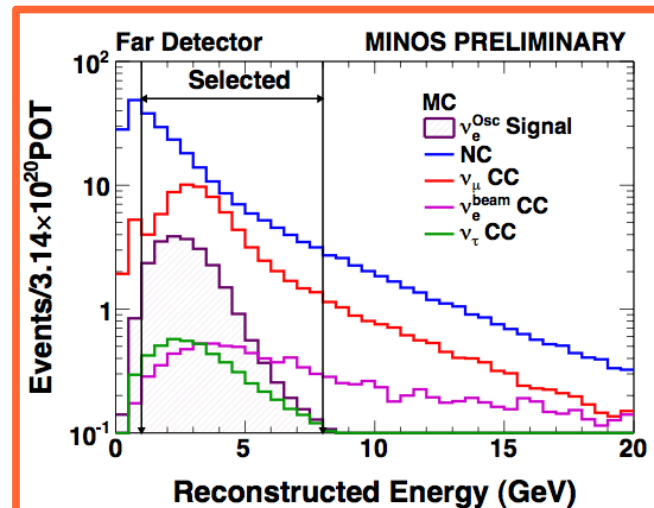
# Preselection



Track length



At least one shower; 4 hit planes in a row



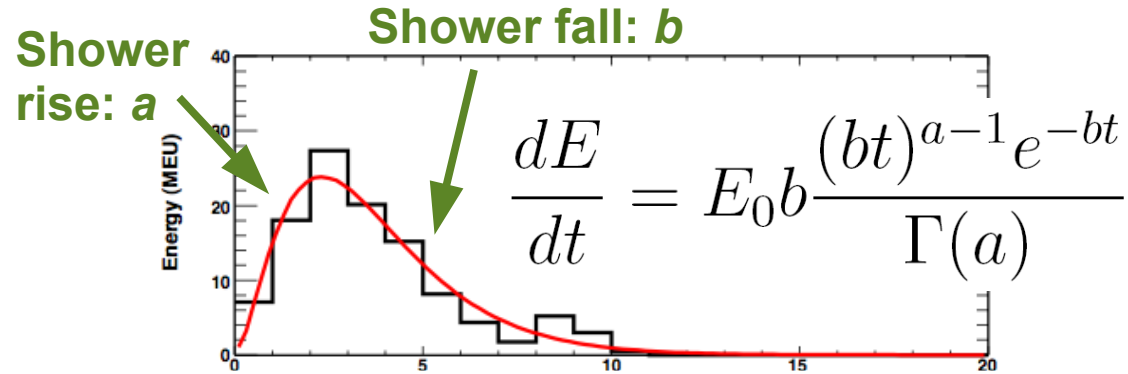
Reconstructed energy



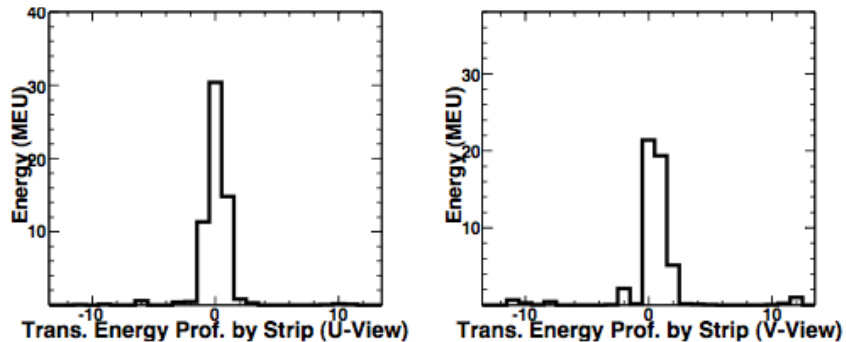
# Data Reduction

	data	eff	MC	eff
fiducial	1188		1221.4	
trk	444	37.4%	445.6	36.5%
trklike	410	34.5%	410.3	33.6%
n shw	406	34.2%	406.4	33.3%
cont. pln	286	24.1%	298.5	24.4%
E > 1 GeV	271	22.8%	282.8	23.2%
E < 8 GeV	227	19.11%	229.3	18.8%

# ANN variables

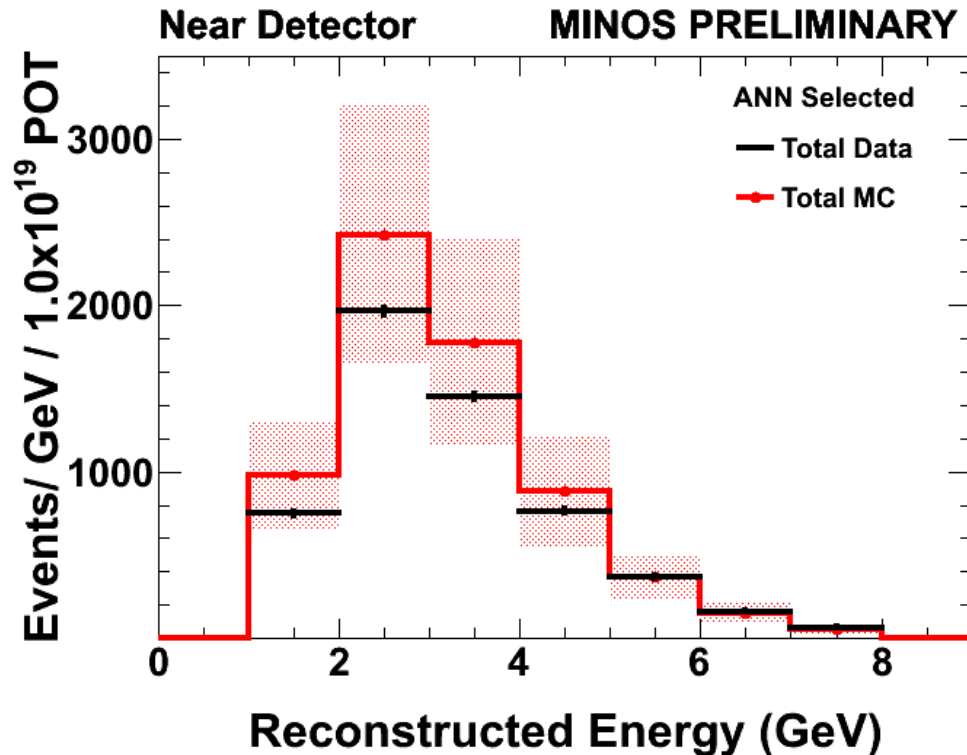


Longitudinal Energy Profile by plane



Transverse Energy Profile by strip

# Near Detector $\nu_e$ candidates



The events that pass our selection cuts are from a kinematic region where hadronization is not well-modeled

→ Large uncertainties on the MC

→ Data/MC discrepancy not surprising

Each background component (CC  $\nu_\mu$ , NC, beam  $\nu_e$ ) must be extrapolated to the far detector separately.

We should not rely on MC for the relative size of CC  $\nu_\mu$ , NC, and beam  $\nu_e$  components of the background.

→ we have two data-driven methods

# Data-Driven Background Separation: Horn On/Horn Off

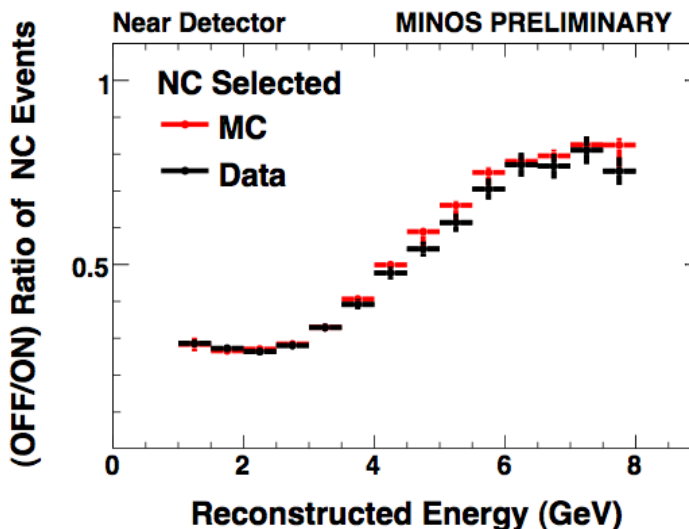
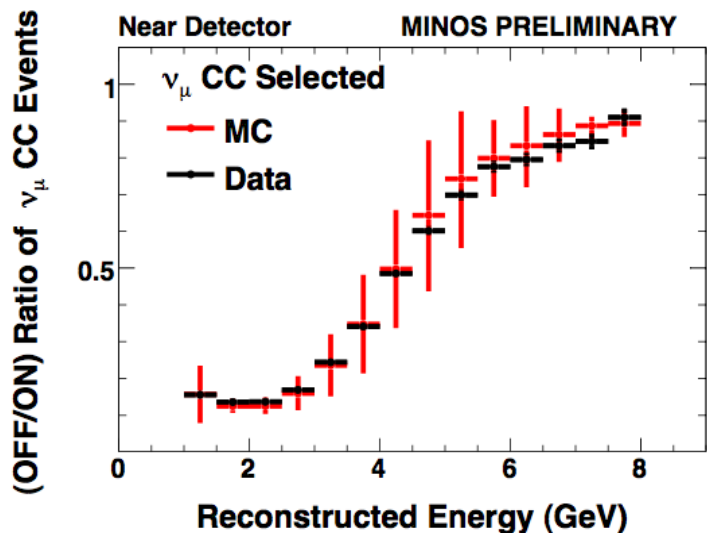
Number of selected data events in the horn off and horn on configurations can be related by the **horn off to horn on ratios** for each component.

Solve for  $N_{CC}^{ON}$  and  $N_{NC}^{ON}$

$$\begin{aligned} \text{data} \rightarrow N^{ON} &= N_{CC}^{ON} + N_{NC}^{ON} + N_e^{ON} \\ N^{OFF} &= r_{CC} N_{CC}^{ON} + r_{NC} N_{NC}^{ON} + N_e^{OFF} \end{aligned}$$

where  $r_x = N_x^{OFF}/N_x^{ON}$

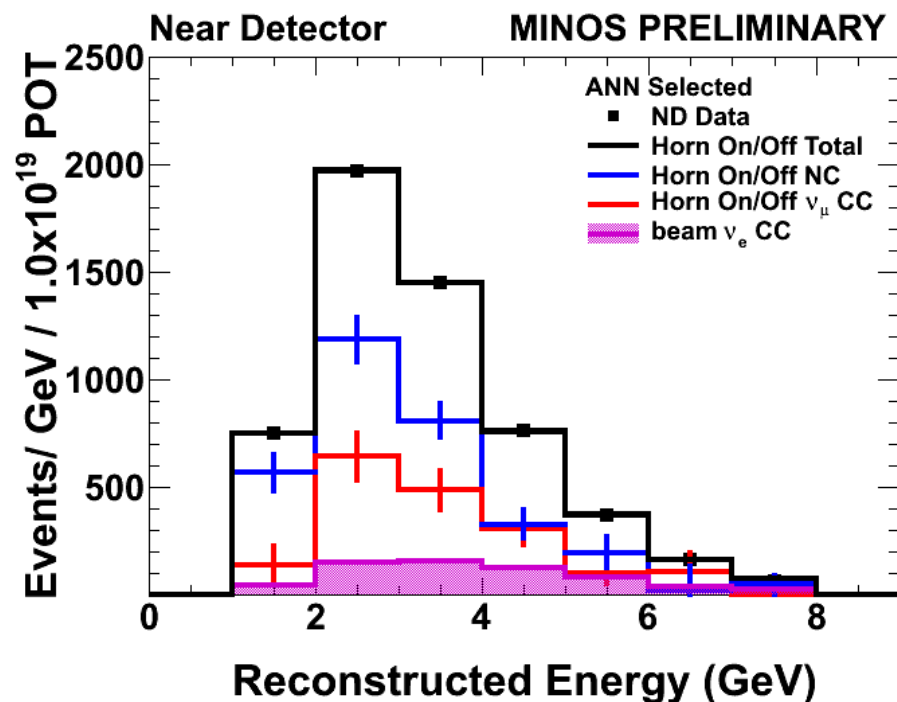
MC (green arrow) points to  $r_x$ . MC (pink arrow) points to  $N_e^{ON}$  and  $N_e^{OFF}$ .



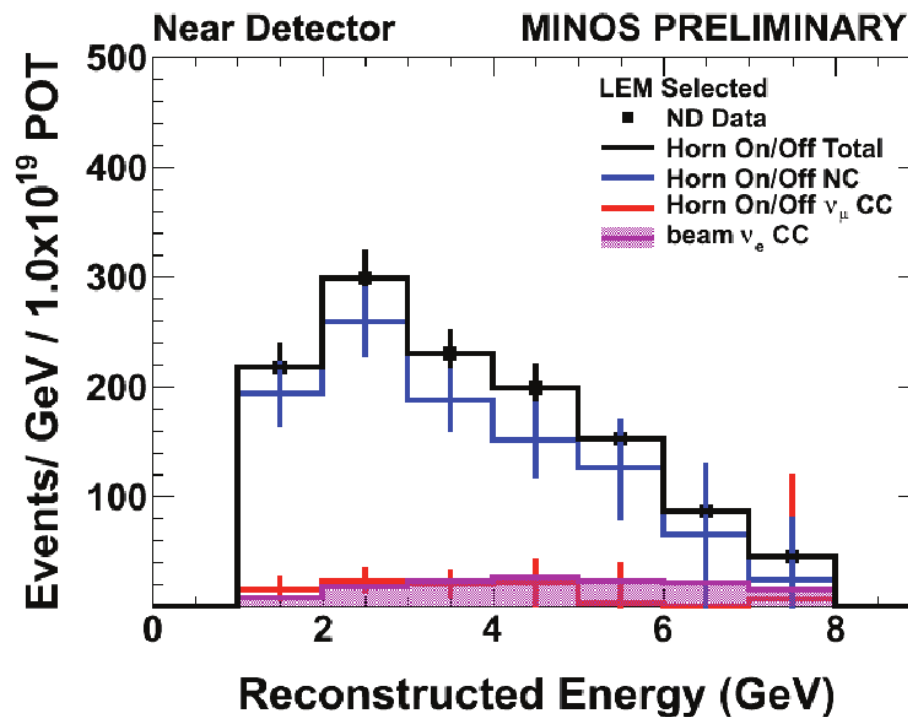
Select CC-like and NC-like events in the fiducial volume and compare data and MC horn off to horn on ratios – **MC models the ratios well.**

# Data-Driven Background Separation: Horn On/Horn Off

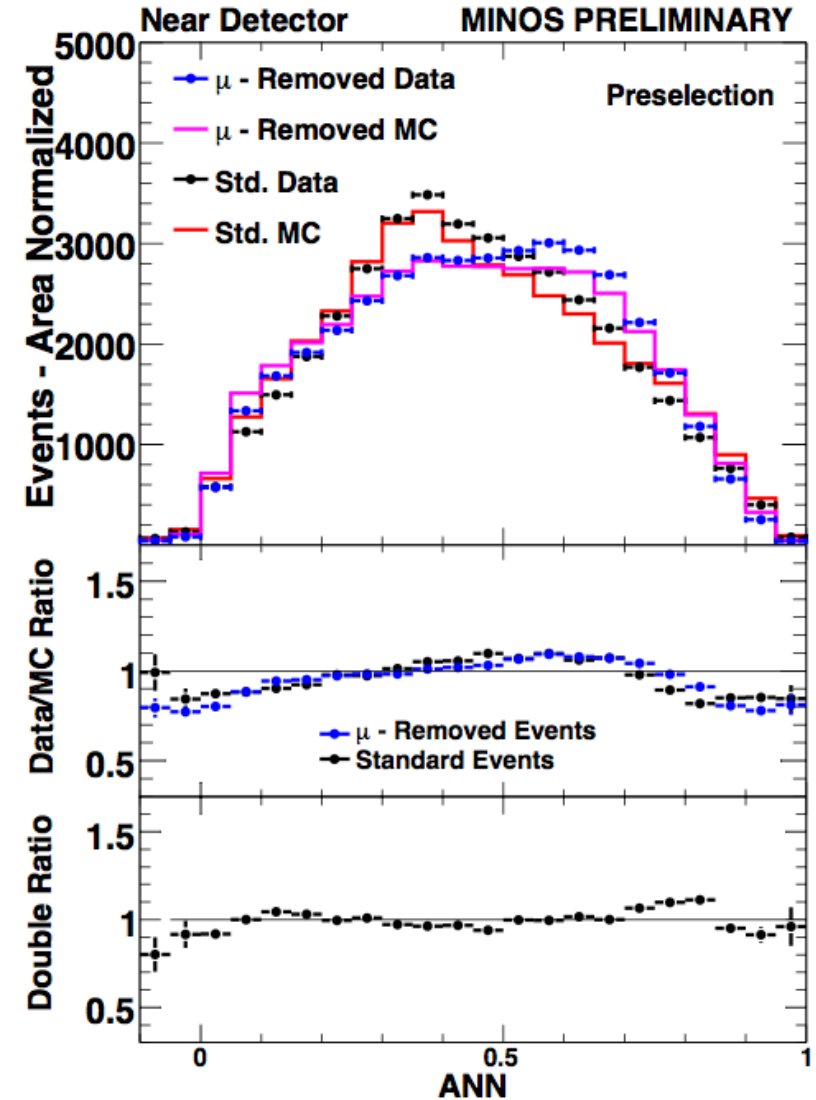
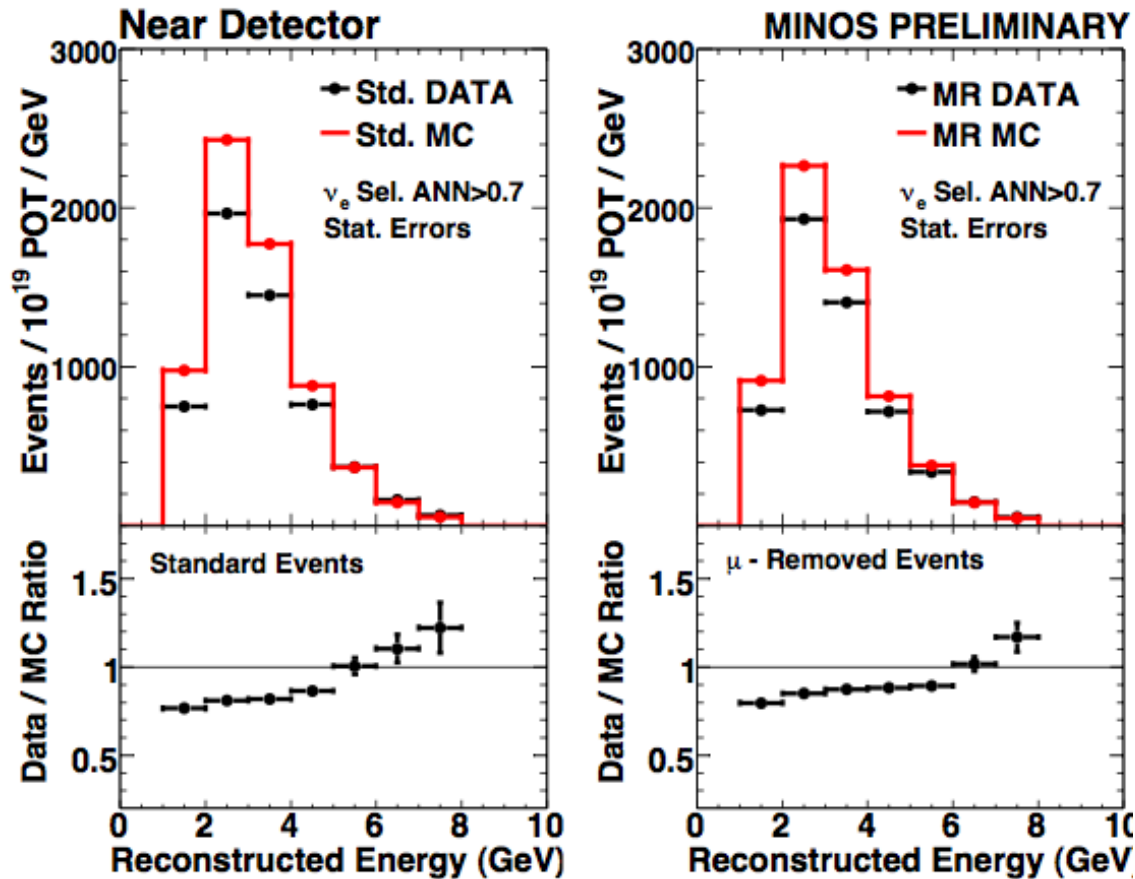
## Horn On Data



## Horn Off Data

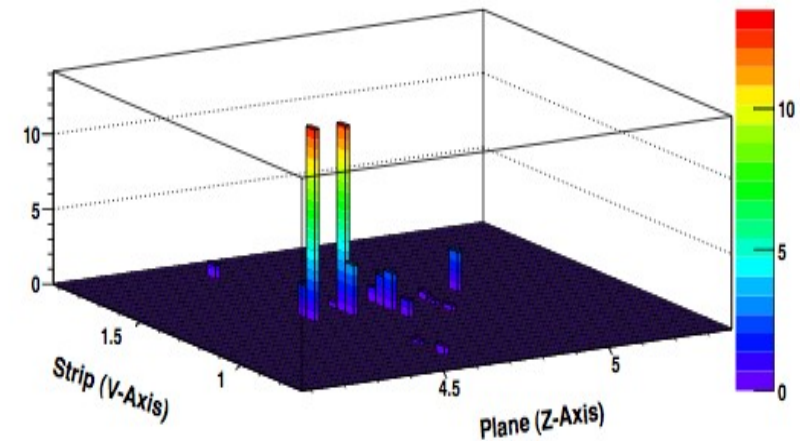
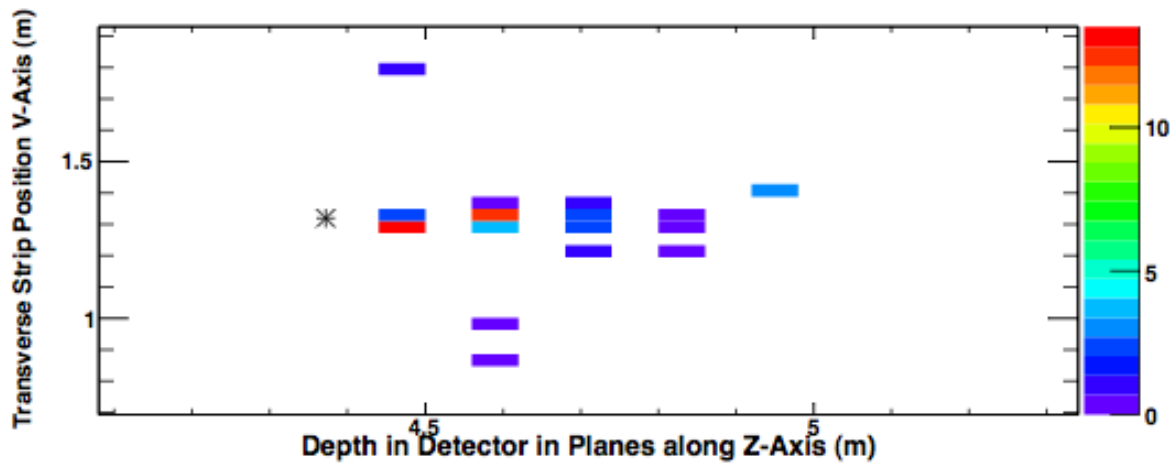
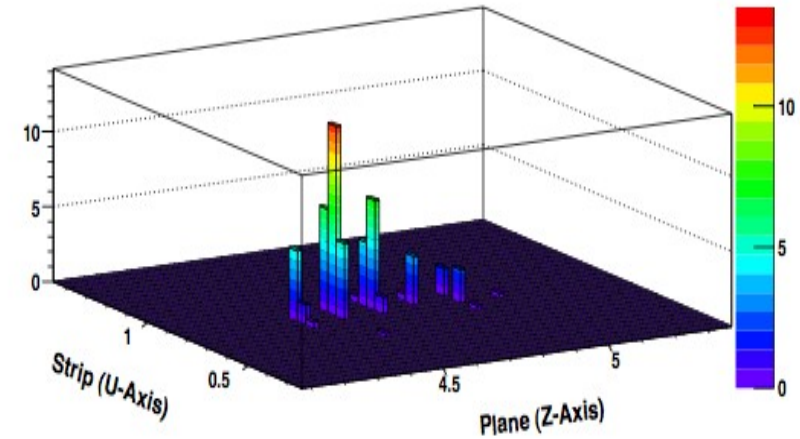
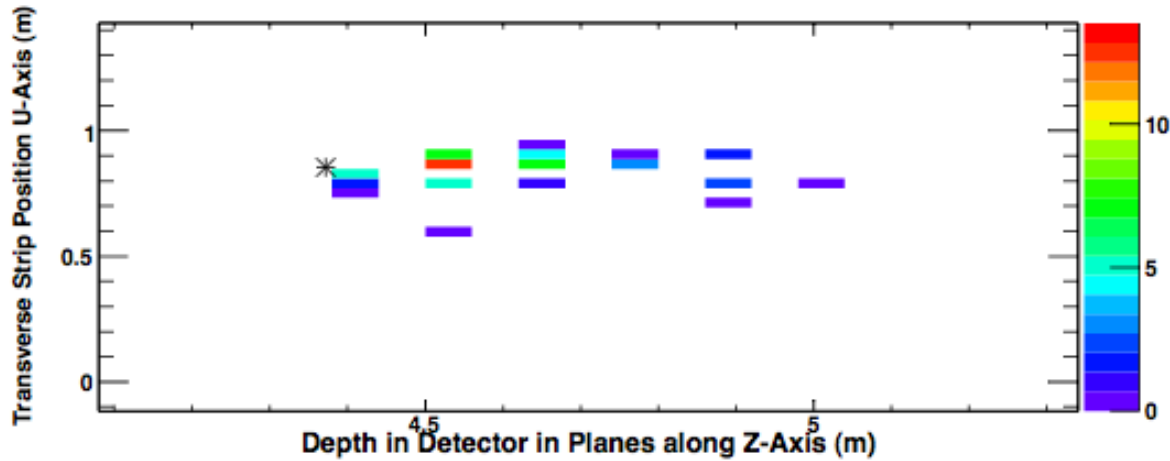


# Data-Driven Background Separation: MRCC



discrepancy between MRCC data and MC is very similar to the discrepancy in standard data and MC, both in shower shape and energy

# Far Detector $\nu_e$ Candidate Event Display



reconstructed energy: 4.6 GeV



# MRE and Signal Efficiency

## Muon Removed w/ Electron Added (MRE)

Take muon removed events, add an electron and re-reconstruct

Allows us to simulate signal events with a real hadronic shower

Apply  $\nu_e$  selection to MRE data and MRE MC – ratio is used to correct signal efficiency

